INTO THE STARLIGHT: The End of the Cosmic Dark Ages

March 3rd - 9th, 2019

firstgalaxies.org

Aspen Center for Physics March 05 2019

Revealing Cosmic Sunrise: Galaxy Build-up in the First 600 Myr

Garth Illingworth University of California Santa Cruz

figure credit: Adolf Schaller

insight into the epoch of the first galaxies — inspired by JWST "first light" goal



reionization epoch – Planck 2016/2018 results





remarkable mission

reionization simulation: Alvarez et al. 2009

Planck 2016/2018

constraints on the reionization history

...Thomson optical depth: $\tau = 0.054 \pm 0.007$

...mid-point redshift at which reionization occurs is found to lie at $z = 7.7 \pm 0.7$

...upper limit to the width of the reionization period of $\Delta z < 2.8$.

...the Universe is ionized at much less than the 10% level at redshifts above $z \simeq 10$... (<1% above $z \simeq 15$)

...an early onset of reionization is strongly disfavored by *Planck* data

Planck Collaboration XLVII + 2016 Planck Results VI Cosmological Parameters + 2018 Planck Results I Overview and Legacy + 2018



....*Planck* data prefer a late and fast transition from a neutral to an ionized universe...

...the Universe is ionized at much less than the 10% level at redshifts above $z \simeq 10$... (<1% above $z \simeq 15$).

....non-standard early galaxies or significantly evolving escape and clumping factors are no longer required.

....nor do the *Planck* results require *any* emission from high-redshift (z = 10-15) galaxies.



we now know when

galaxies really began to reionize the universe (at z<10)

this is a crucial piece of information re the epoch of the first galaxies

reionization history compared with observational astrophysical constraints



figure from Planck Collaboration XLVII + 2016

Bouwens+2015 Robertson+2015 Ishigaki+2015

striking consistency with Hubble results

indicates that galaxies were responsible for reionization

reionization history compared with observational astrophysical constraints



when did the "first stars" appear?



EDGES:

Experiment to Detect the Global Epoch of Reionization Signature

EDGES low-1 antenna Murchison Radio-Astronomy Observatory (MRO) Tet.

timing the "first stars"

when the "first stars" started to produce UV L α photons



Age of the Universe (Myr)

EDGES: first stars become prominent at redshift z~20 (~180 million years)

timing the "first stars"



EDGES: first stars become prominent at redshift z~20 (~180 million years)

what constraints do we have on the first galaxies?

searching for the earliest galaxies





Oesch + 2014

Hubble

Spitzer



Oesch + 2014, 2016

Hubble

the most distant galaxy found to date

surprising discovery of GN-z11



detection of GN-z11 in *existing data* is unexpected, given current models

Oesch+2016

gdi

GN-z11

simulations show that galaxies as massive as GNz-11 at z~11 are rare but not unexpected *per se*



GN-z11 observed properties

mass $10^9 M_{\odot}$ SFR $24 M_{\odot}/yr$ $A_{UV} < 0.2 \text{ mag}$ age 40 Myr β -2.5



Mutch+2016

GN-z11



at z~11 are rare but not unexpected *per se* GN-z11 observed properties

simulations show that galaxies as massive as GNz-11

mass 10^9 M_{\odot} SFR 24 M $_{\odot}$ /yr β -2.5 A_{UV} <0.2 mag age 40 Myr

the derived physical properties of GNz11 are consistent with expectations from large-volume simulations



Mutch+2016

but it is unexpected to find GN-z11 in such small search volumes/areas (by factor 10-100)?

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DRAGONS



Spitzer IRAC

Planck Collaboration Results I + 2018



Spitzer IRAC

Planck Collaboration Results I + 2018

what constraints do we have on the first galaxies?

searching for evidence of early star formation

MACS1149–JD1

z=9.11 from ALMA [OIII]

Hashimoto et al (and Katz et al) show large IRAC break in JD1 SED interpreted as a Balmer break, indicating old population that formed 300 Myr earlier at z~15

IRAC color

from

Zheng et al 2017

very puzzling since Balmer breaks are quite rare at somewhat lower redshifts – where samples are large

errors on the IRAC photometry are surprisingly small





single panel from Fig 9 – Katz+2019

Hashimoto+2018



what constraints can Hubble and Spitzer put on the first galaxy epoch?

what do the highest redshift galaxies at z~10 (480 Myr) tell us

z~10 (500 Myr) galaxies are hard to find!



from 8 years of WFC3/IR imaging

searched every WFC3/IR dataset but we find only 9 z~10 galaxies (at ~500 Myr)

gali

Oesch+2017

see also: Zheng+2012; Coe+2013; Bouwens+2013,15,16; Ellis+2013; McLure+2013; Ishigaki+2014,17; Infante+2015; Bernard+2016; Calvi+2016; McLeod+2016

model comparisons – the luminosity function at z~10



considerable spread shape matches (broadly) to models – but models are consistently high

"accelerated evolution" – the star formation rate density at z~9-10

clearly a trend to lower SFRD at z>8 than initially expected

"accelerated evolution" is actually consistent with the expected buildup* of dark matter halos over that time

*dark matter halo growth (>~ $10^{10} M_{\odot}$) from HMFcalc – Murray+2013



Oesch+2013,2014,2017

(Cela

"accelerated evolution" – the star formation rate density at z^9-10

clearly a trend to lower SFRD at z>8 than initially expected

"accelerated evolution" is actually consistent with the expected buildup* of dark matter halos over that time

Note: this result also indicates that the star formation efficiency (SFE) does not evolve significantly with cosmic time at z>6



Oesch+2013,2014,2017

<u>a</u>h

model comparisons – the star formation rate density at z>6

note that there is a large range of shapes/slopes!







way fewer galaxies than expected at redshift 10!



way fewer galaxies than expected at redshift 10!



can JWST find the first galaxies?

will they be so rare that they will be hard to find?

will they occur at such high redshifts that they will be hard for JWST to see?



galaxies evolving rapidly at z>9 + reionization turn-on at z~10 (<1% at z~15)

suggest major changes in galaxy population at z~10-12-15

great for JWST's "first light" goal since galaxies likely evolving rapidly at z~10-12-15 and JWST will be able to detect galaxies out to z~15










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James Webb Space Telescope JWST

JWST STATUS UPDATE

NASA ESA CSA

NGST started at STScI in the mid-late 1980s by Pierre Bely, Garth Illingworth and Peter Stockman

THE NEXT GENERATION SPACE TELESCOPE



Proceedings of a Workshop held at the Space Telescope Science Institute Baltimore, Maryland, 13-15 September 1989

1989





Riccardo Giacconi encouraged us in 1986/7 – "start early on the next mission" – he told us that it would take a long time!

ASTRONOMY AND ASTROPHYSICS

THE DECADE OF DISCOVERY IN .

NATIONAL RESEARCH COUNCIL

WORKING PAPERS

Astronomy and Astrophysics Panel Reports

1991

NATIONAL RESEARCH COUNCIL

ASTROTECH 21 WORKSHOPS SERIES II

VOLUME

4

1991

SERIES II MISSION CONCEPTS AND TECHNOLOGY REQUIREMENTS

Workshop Proceedings: Technologies for Large Filled-Aperture Telescopes in Space





September 15, 1991

JPL D-8541, Vol. 4

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THE NEXT GENERATION SPACE TELESCOPE

Simulated images of NGC2903 translated to Z=1

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JEL 0-0541, Vol. 4

1990 Decadal Survey: UV-Optical in Space Panel recommended:

- 6-m passively-cooled infrared telescope
- *derived a cost of \$2B in FY90\$ (~\$4B in 2018\$)*
 - for launch in 2009 to a high orbit

OTIS in JSC Chamber A in mid-2017 for the three month cryogenic vacuum test

and -

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JWST flight sunshield in Northrop Grumman clean room

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INT THE CASE









JWST launching on the European Ariane 5 in about 2 years





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reionization constraints from Planck 2018



GN-z11 is a pathfinder into the epoch of the earliest galaxies

Planck Collaboration



Oesch + 2014, 2016

Hubble

Spitzer

JWST sunshield in Northrop Grumman clean room

TALL TALLAS



James Webb Space Telescope JWST

operational in 2021

NASA ESA CSA

our "fírst galaxy epoch" telescopes









what constraints do we have on the first galaxies?

searching for the earliest galaxies

MACS1149–JD1

z=9.11 from ALMA [OIII]

comparison of IRAC measurements from three different groups

Hashimoto+ & Katz+ IRAC color 3.6-4.5 = 0.91 ± 0.18 ** this is ~5 σ ** Shipley+ IRAC color 3.6-4.5 \sim 0.3±0.? ** this is $\sim 1\sigma$ ** our recent IRAC color 3.6-4.5 = 0.51 ± 0.26 ** this is $\sim 2\sigma$ **









the inconsistencies in the IRAC photometry need to be resolved

vertical scales are ~same – horizontal differ