High precision wide-field instrument and foreground simulations for EoR experiments

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The Foreground Problem



Parsons et al. (2012)

Fourier Space and Delay Spectrum



Parsons et al. (2012)

Motivation for High Precision Modeling



- >10-sigma statistical detection expected with ~1000 hours data
- Currently heavily limited by foregrounds and instrument systematics (e.g. PAPER64 Ali et al. 2015, Pober et al. 2015; MWA Dillon et al. 2013)

Precision Radio Interferometry Simulations (PRISim)

- **Objectives with PRISim:**
- Comprehensive all-sky simulations (with good match to data)
- Role of Wide-field measurements

- Role of compact, diffuse foregrounds
- Role of instrument such as antenna aperture and its chromaticity
- Solutions to mitigate systematics

Model-Data Agree well



Impact of diffuse, compact emission with LST



Diffuse Emission

Point sources

Mitigation of systematics via Antenna Geometry



HERA HI/FG Sensitivity vs. Beam Chromaticity



Uniform Disk Airy Pattern –

Simulated Chromatic HERA beam

- Differences seen only due to spectral differences in Antenna beam
- Beam chromaticity worsens foreground contamination
- HERA is sensitive to EoR nevertheless

Design Specs on Reflections in Instrument



Thyagarajan et al. (2016), under HERA collaboration review

Reflections are inevitable in electrical systems **Reflections extend** foregrounds and contamination in delay spectrum **Require reflected** foregrounds to be below HI signal levels • HERA will beat these specs comfortably

EoR Observing Window Efficiency



150 MHz subband (z=8.47)

170 MHz subband (z=7.36)

- All HERA baselines sensitive to EoR for most of observing window
- Robust to different models and redshifts
- HERA has extreme control over instrumental systematics and foreground contamination

Summary

- PRISim high precision simulations for wide-field radio interferometry publicly available (<u>https://github.com/nithyanandan/PRISim</u>)
- Discovery of new instrument + foreground physics:
- Foregrounds through the instrument are not smooth
- Wide-field effects lead to pitchfork effect diffuse emission near horizon even on long baselines

- Contamination significant from far away from primary field of view due to small but non-zero beam response
- Antenna beam chromaticity and reflections worsen contamination
- Solutions to tackle systematics and the way forward for HERA and SKA-low:
 - Critical to explore antenna apertures and spectral features in future designs
 - Baseline weighting technique prospective for power spectrum estimation methods
 - HERA design robust to systematics offers great promise for EoR detection