



Dan Coe

STScI ESA/AURA Astronomer

ACS & NIRCам Instrument Scientist

Abell 2163
archival ACS imaging
reduced by
Roberto Avlia and Sara Ogaz

Long history of gravitational lensing

see review by Kneib & Natarajan 2011

CFHT (1985)

WFPC2 (1995)

ACS (2009)

ACS + WFC3/IR (2016)
(with more observations to come)



Abell 370

Large Hubble Cluster Lensing Programs



CLASH

Cluster Lensing And Supernova survey with Hubble

PI Marc Postman

25 clusters

524 Hubble orbits

HST Cycles 18 – 20

Nov. 2010 – July 2013

1. cluster science
2. high-z galaxies



Frontier Fields

Directors' Discretionary Time

PI Jennifer Lotz

6 clusters

840 Hubble orbits

HST Cycles 21 – 23

Oct. 2013 – Sept. 2016

faint high-z galaxies



RELICS

Reionization Lensing Cluster Survey

PI Dan Coe

41 clusters

190 Hubble orbits

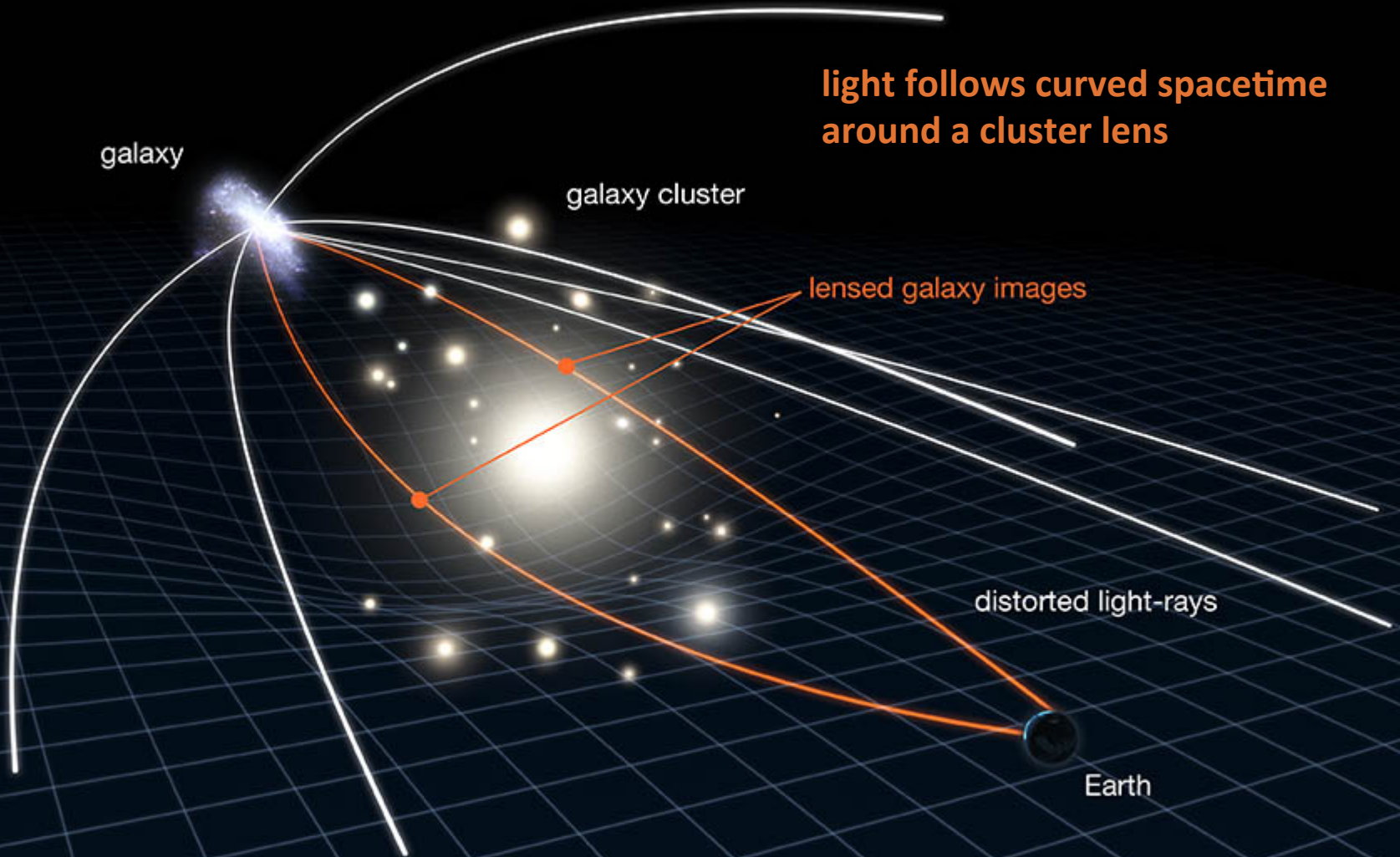
HST Cycle 23

Oct. 2015 – April 2017

bright high-z galaxies

Also GLASS, SGAS, LoCuSS, Ebeling, Kochanek...

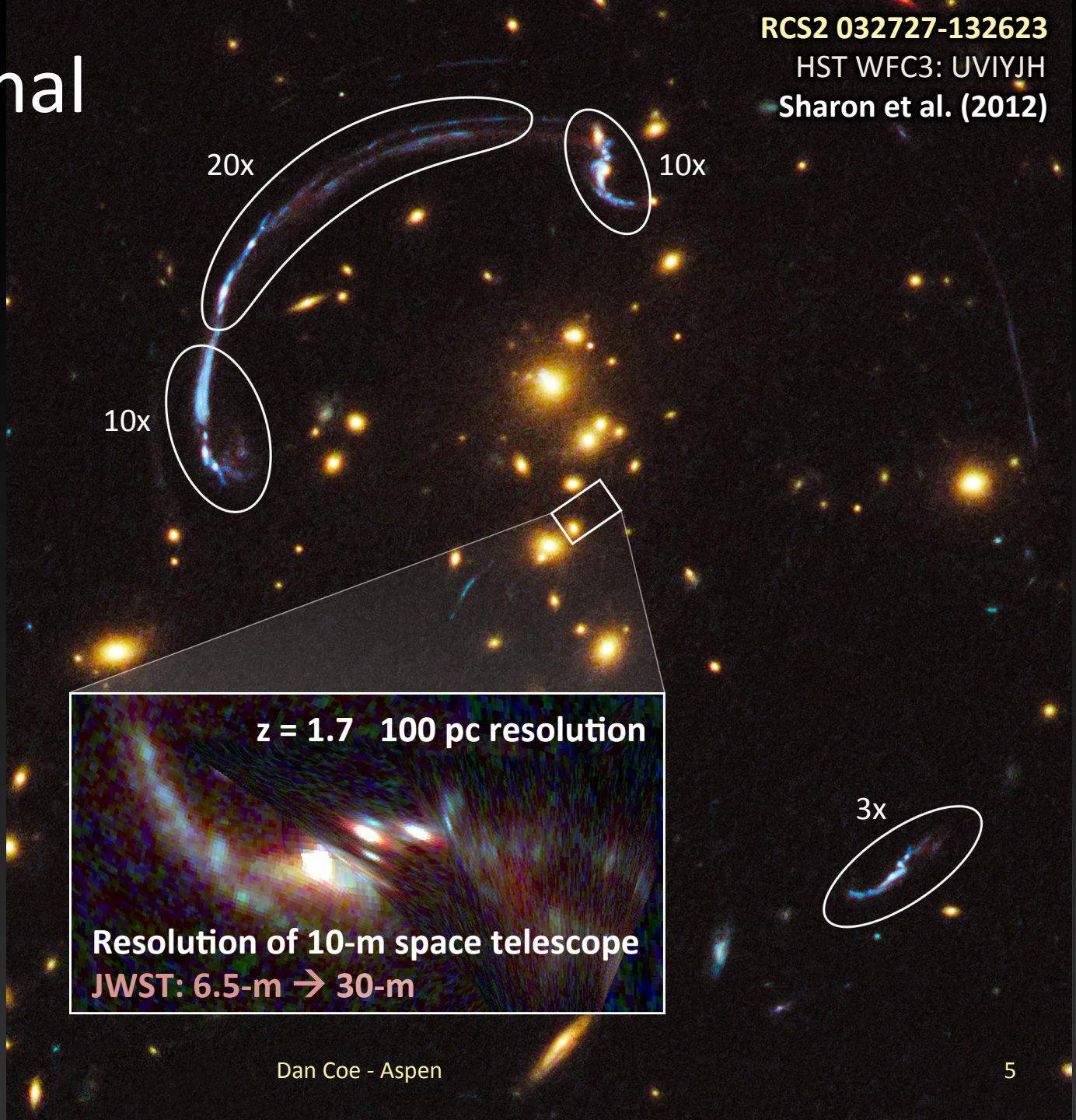
Gravitational lensing magnifies distant galaxies enabling more efficient discovery and detailed analysis



Animation: <http://www.spacetelescope.org/videos/heic1106a/> – NASA, ESA & L. Calçada

Gravitational lensing magnifies the distant universe

RCS2 032727-132623
HST WFC3: UVIIJH
Sharon et al. (2012)

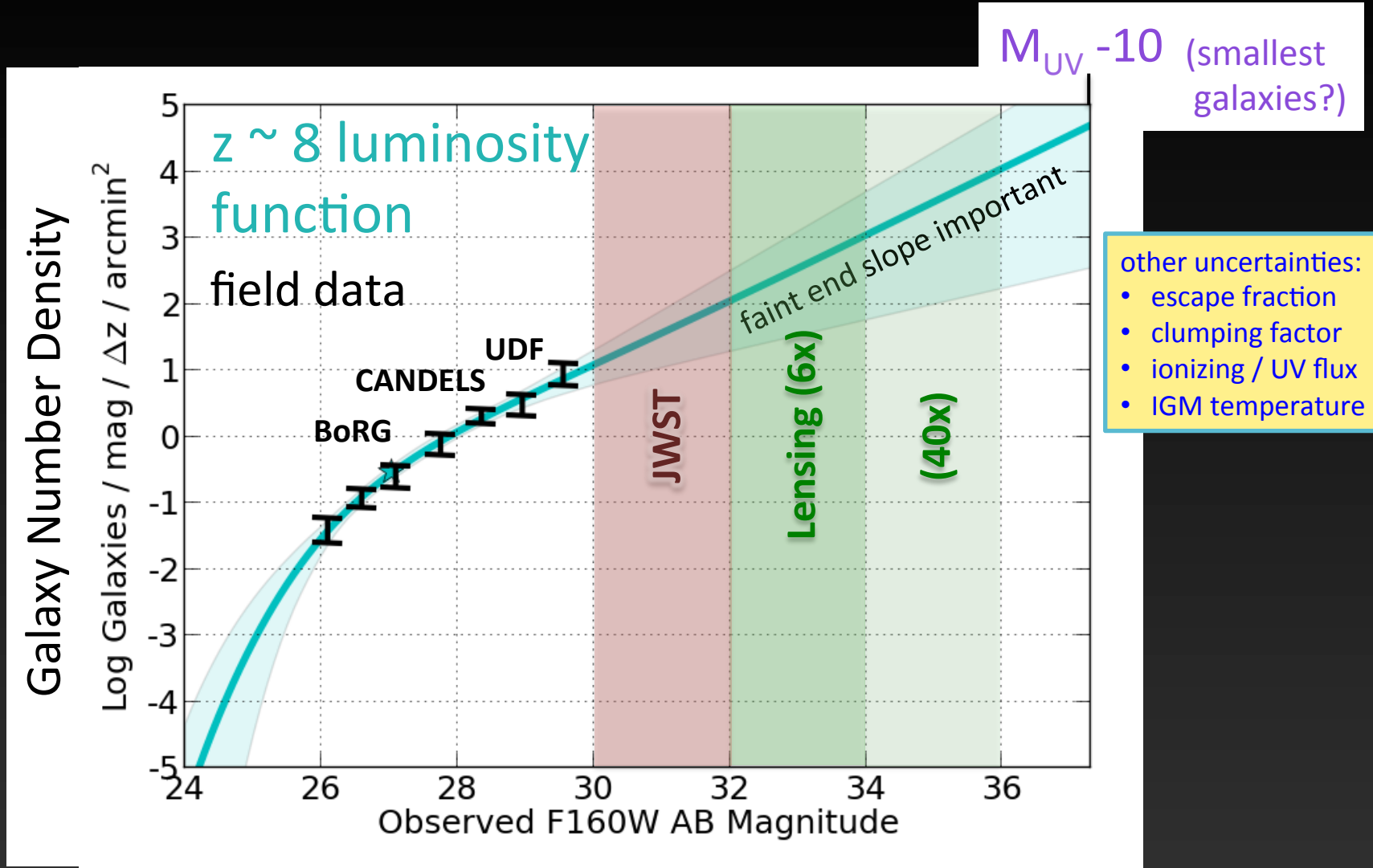


20x
10x
10x

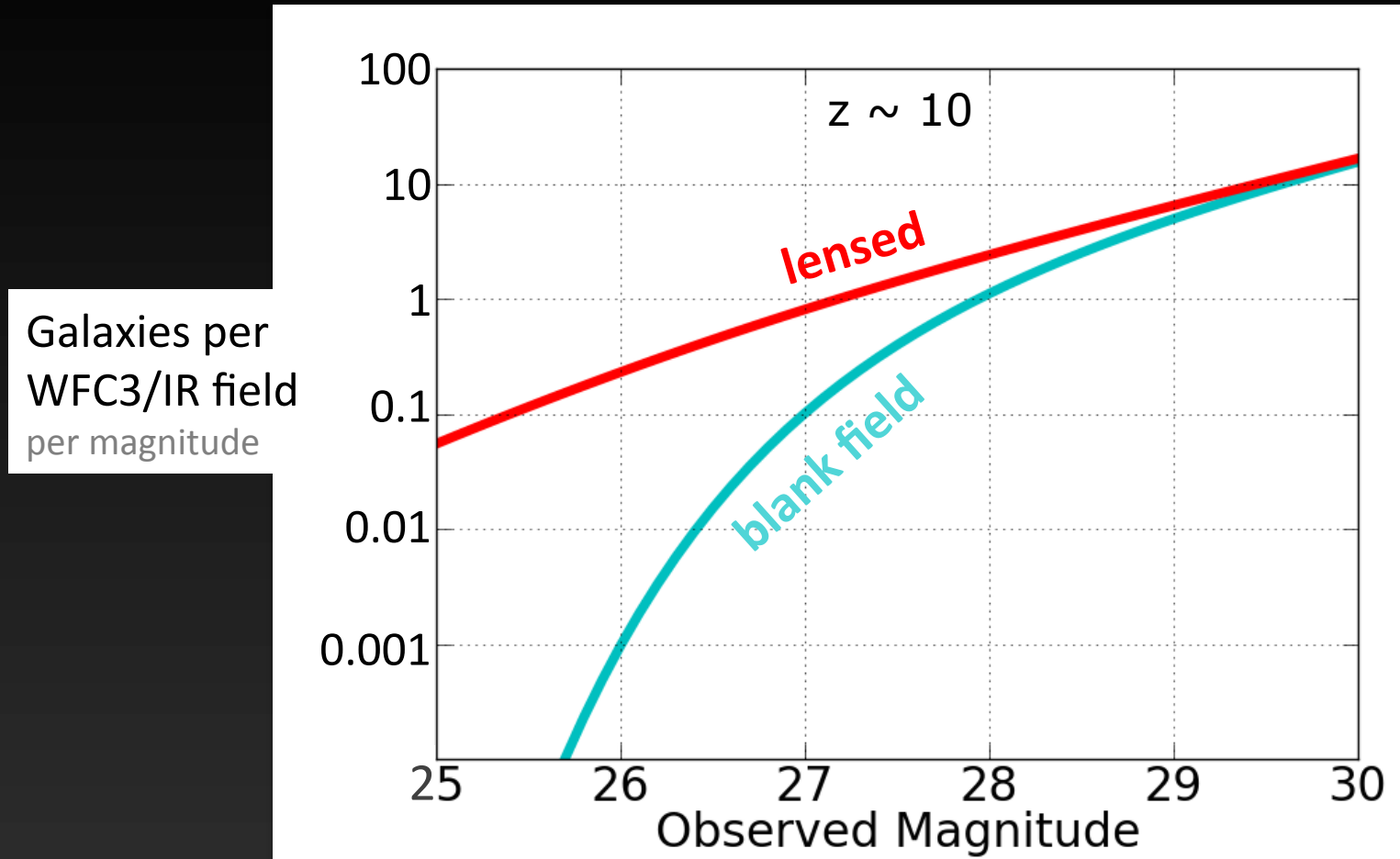
$z = 1.7$ 100 pc resolution
Resolution of 10-m space telescope
JWST: 6.5-m \rightarrow 30-m

3x

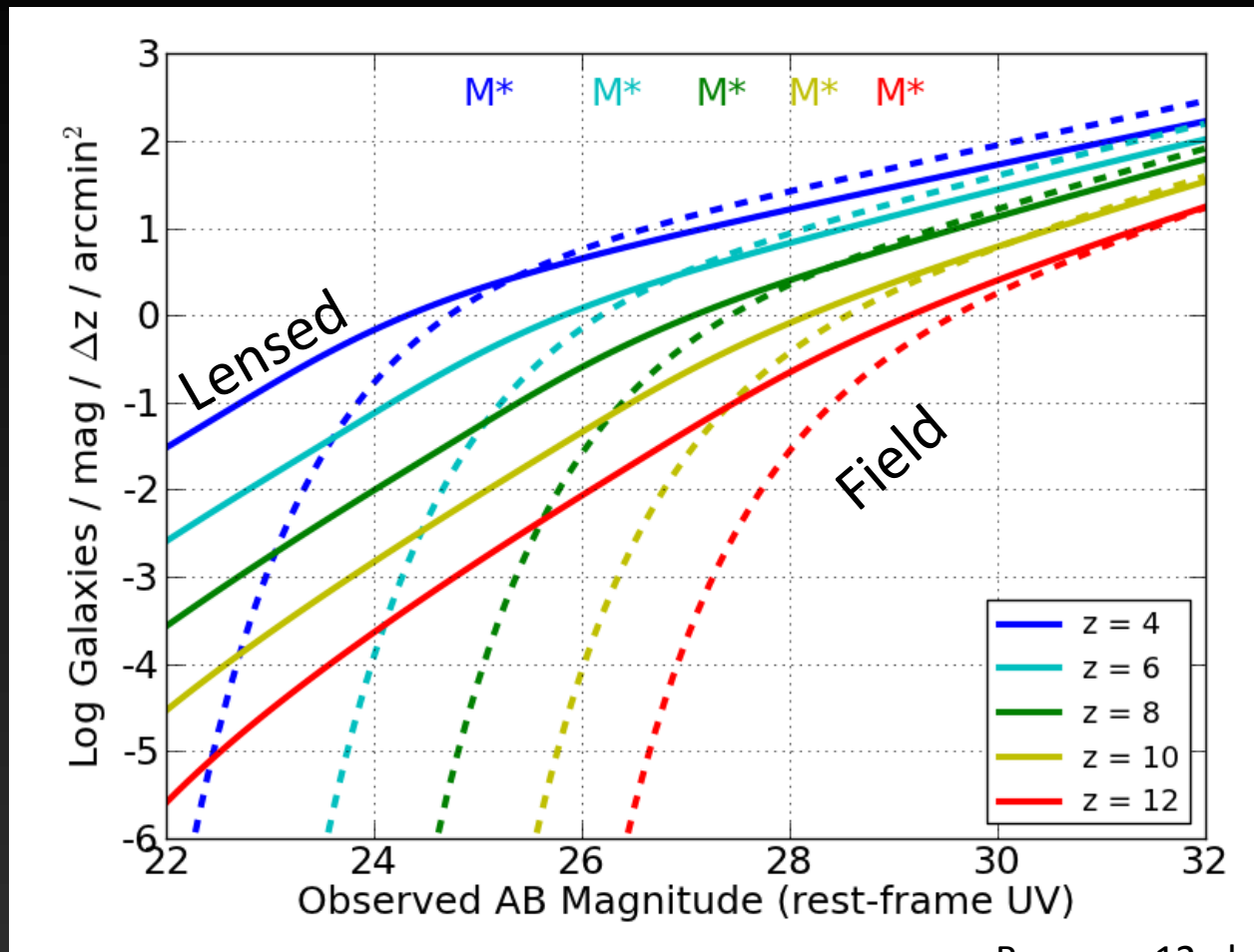
Large extrapolations required to estimate faint galaxy contribution to reionization



Cluster lensing enables more efficient discovery of the most distant galaxies

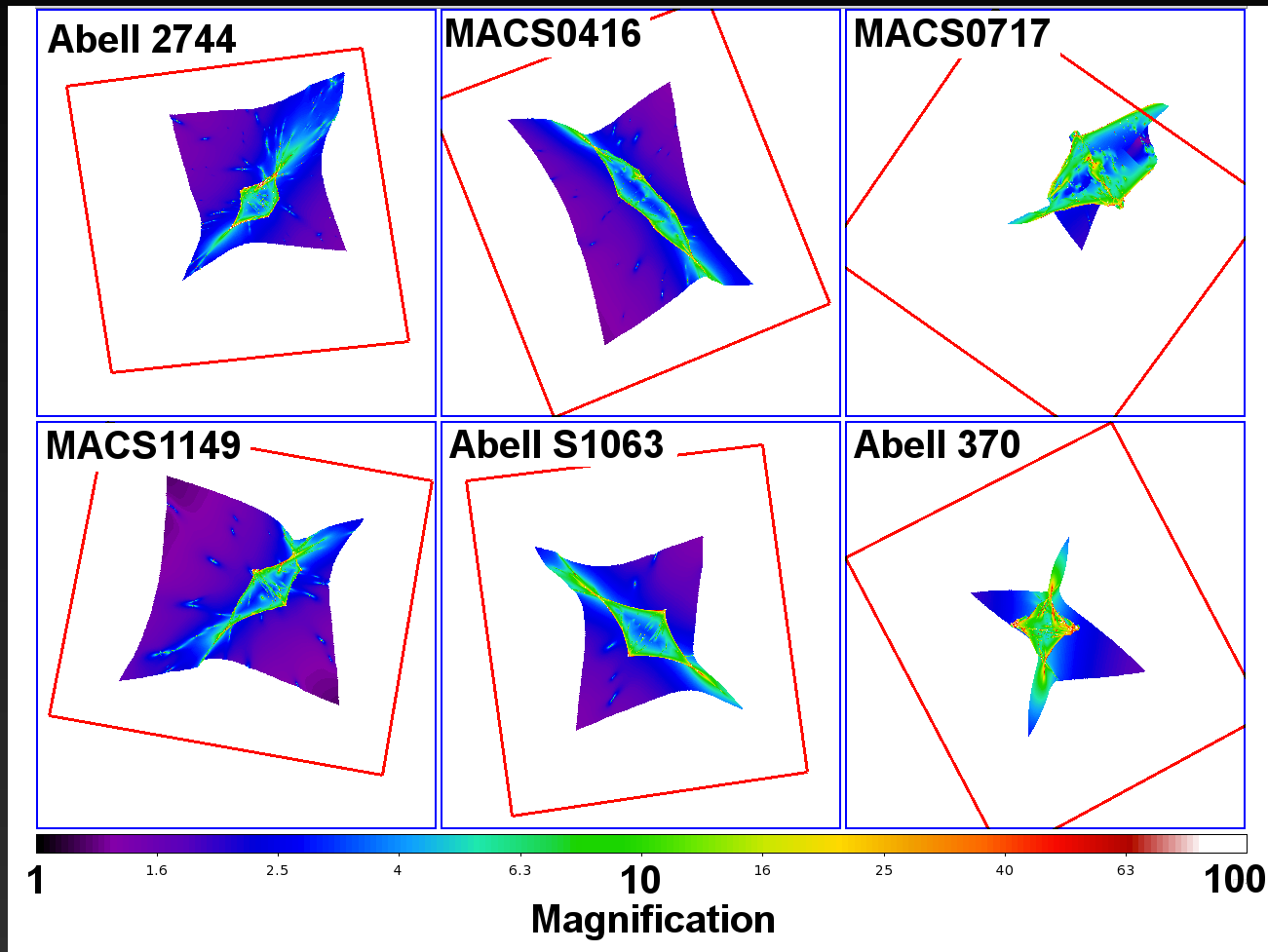


Cluster lensing wins at brighter magnitudes / higher redshifts



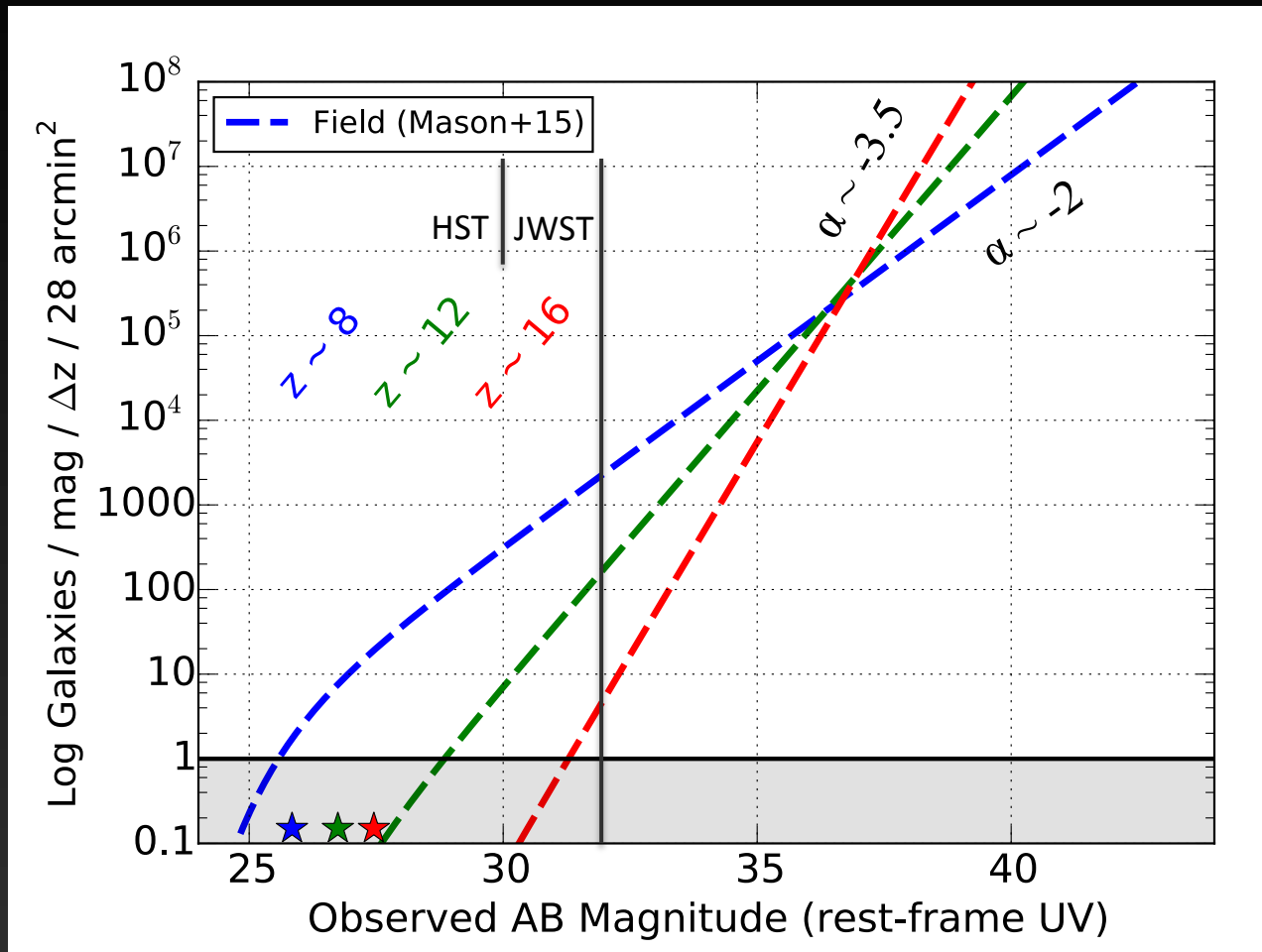
Bouwens12a luminosity function
evolving all 3 parameters

Cluster lensing trades search area for high magnifications

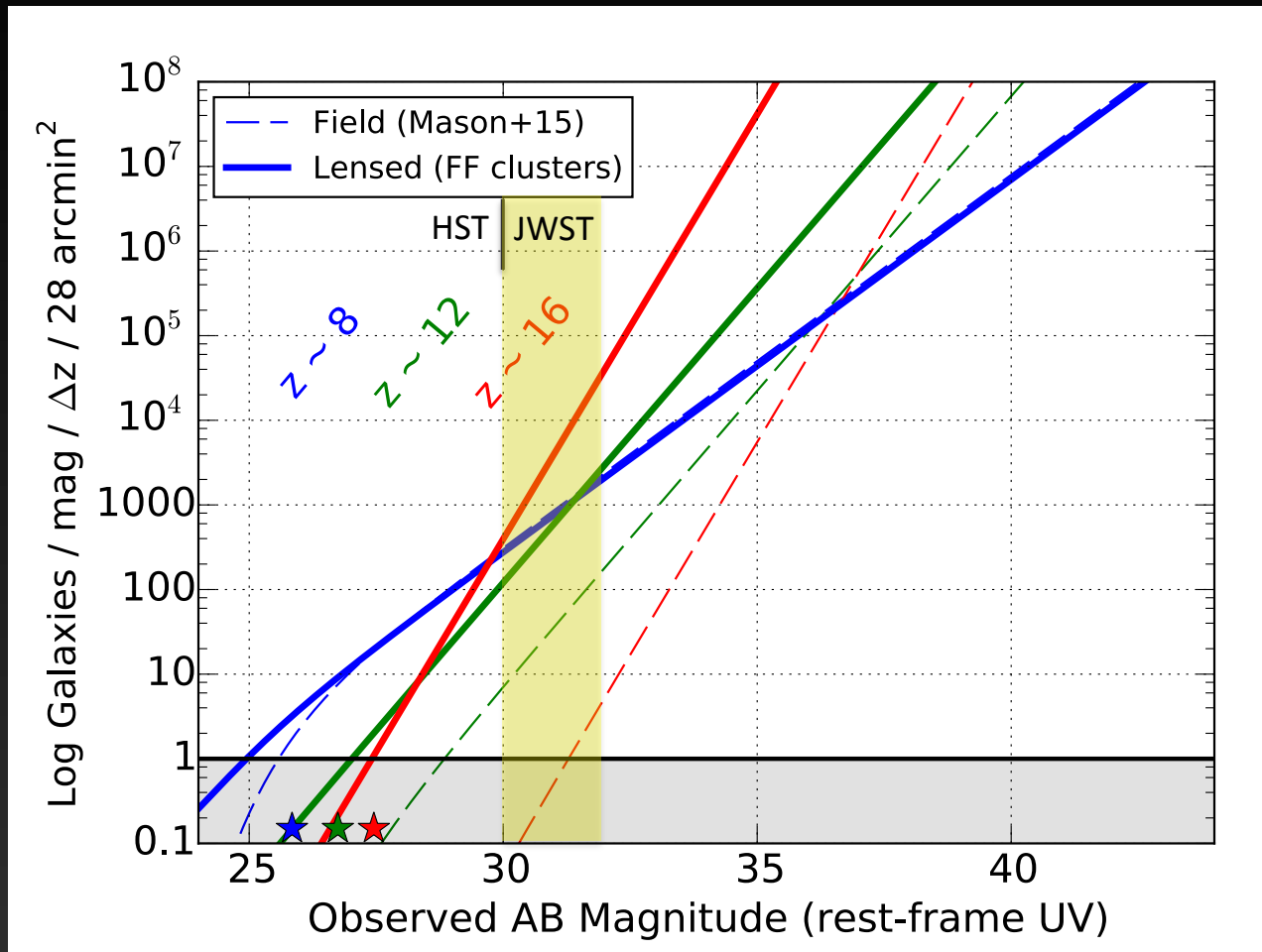


“CATS” team
lens models
of the
Frontier Fields
galaxy clusters

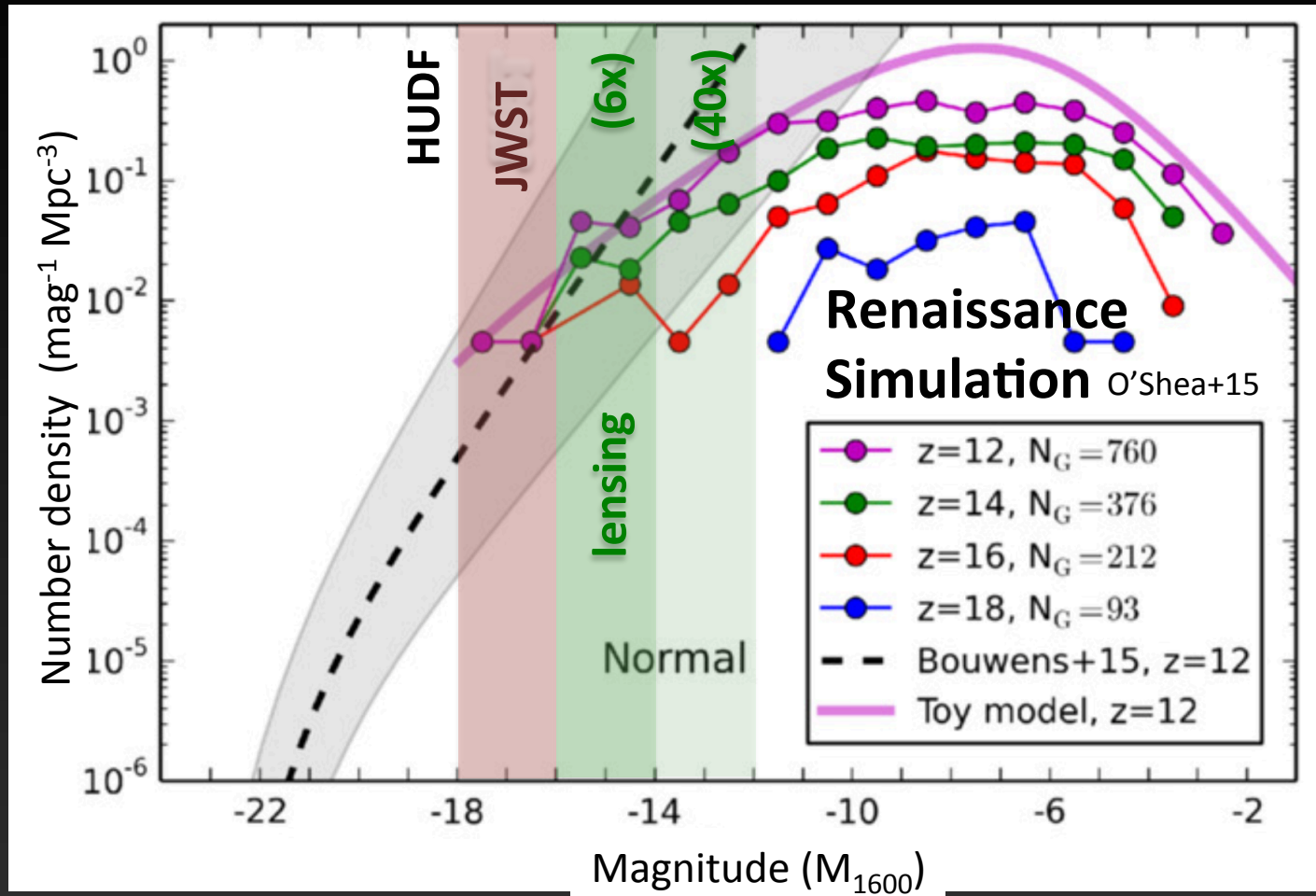
At even higher redshifts, steeper faint end slopes expected



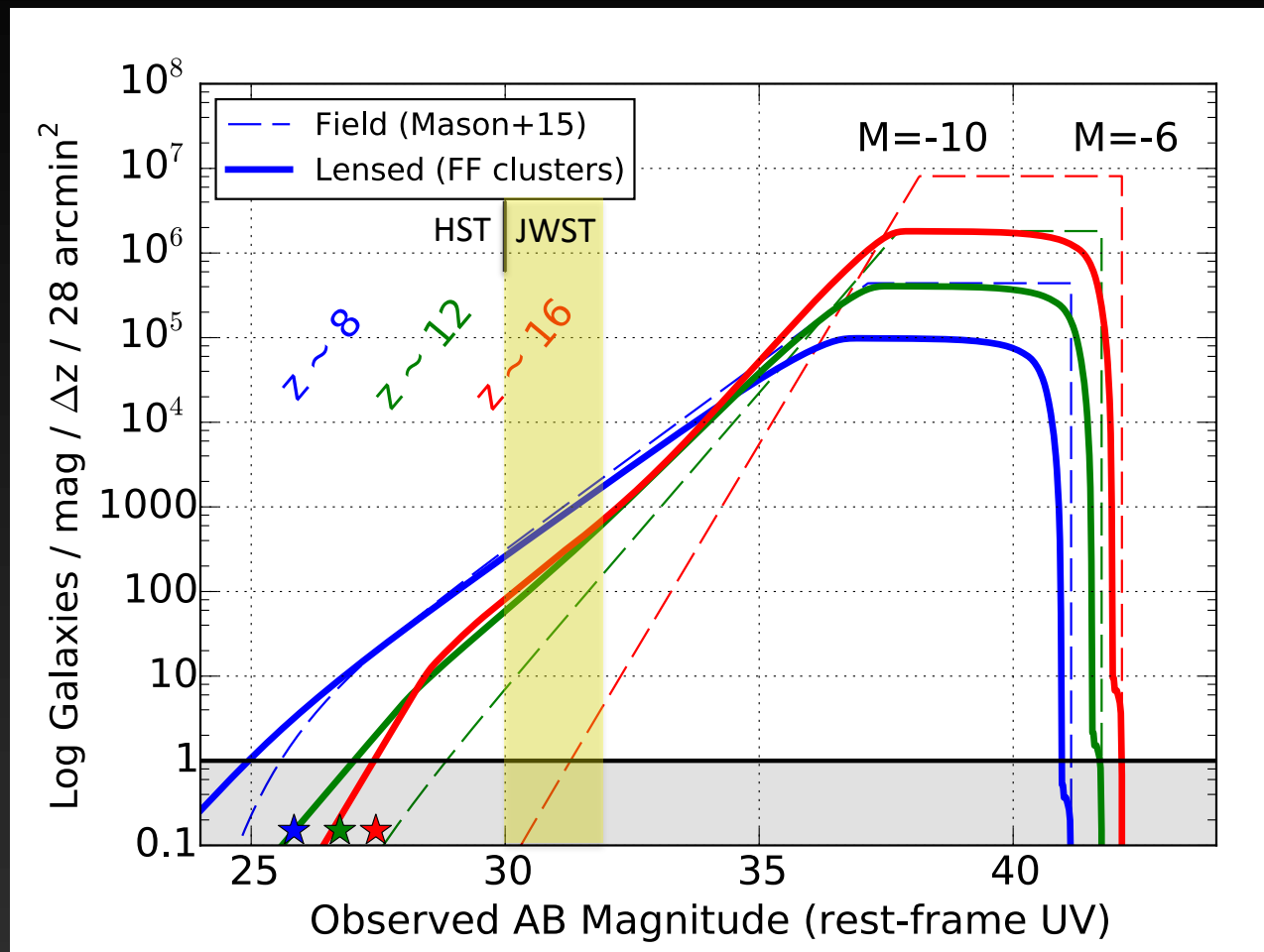
Lensing advantage increases at $z > 10$ thanks to steep faint end slopes



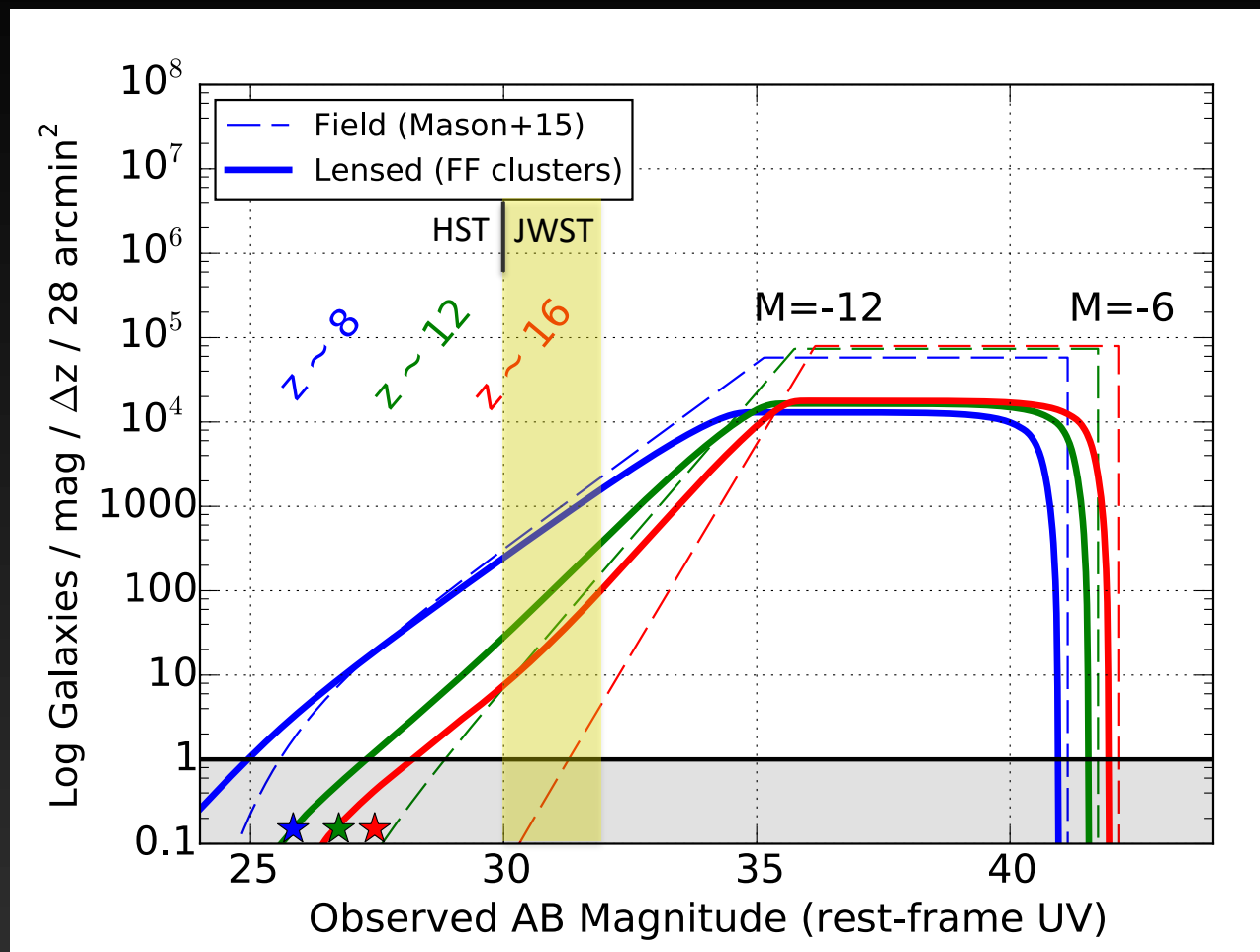
Faint end flattening predicted



Lensing advantage increases at $z > 10$ even if LF is truncated



Lensing advantage increases at $z > 10$ even if LF is truncated

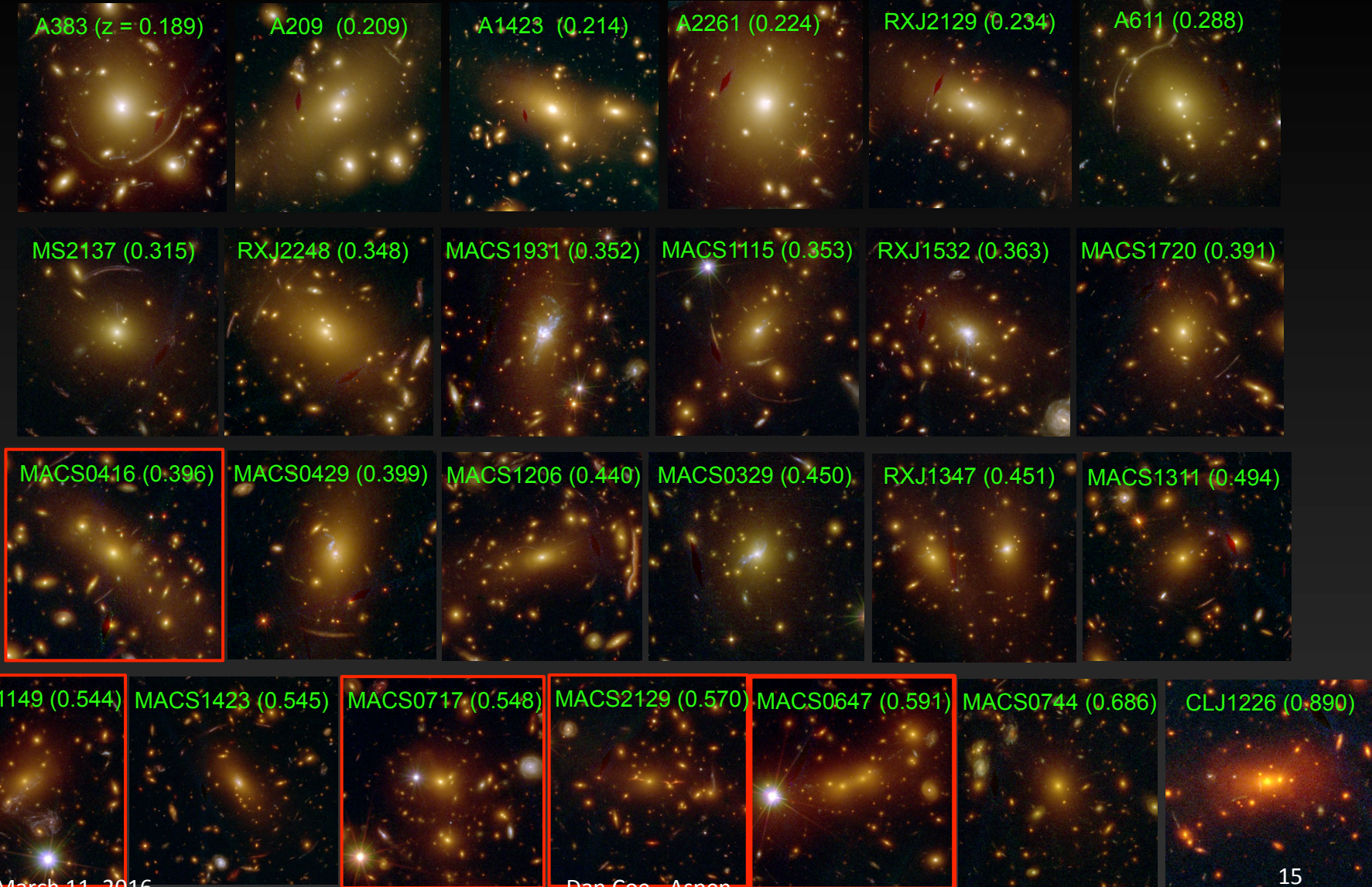




HST observations of 25 clusters

each 20 orbits, 16 filters

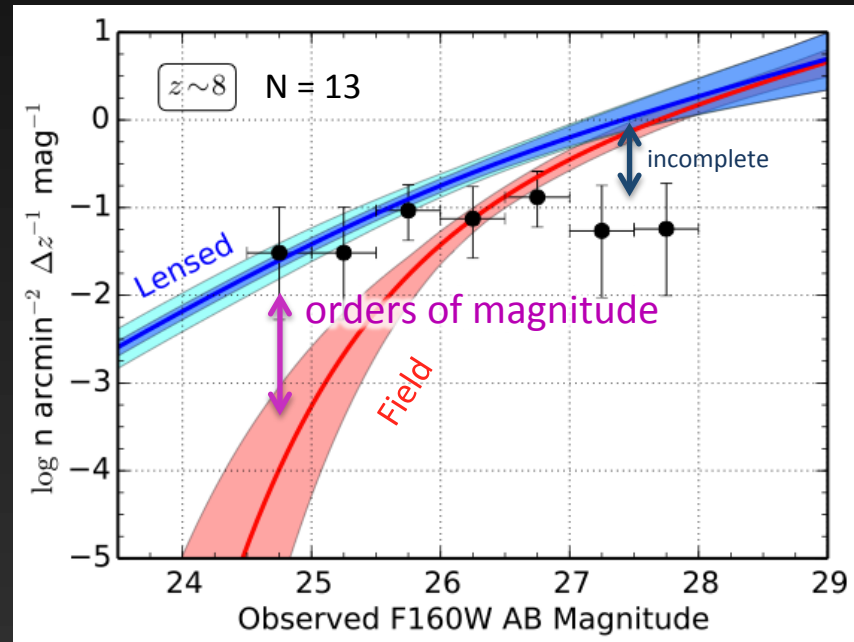
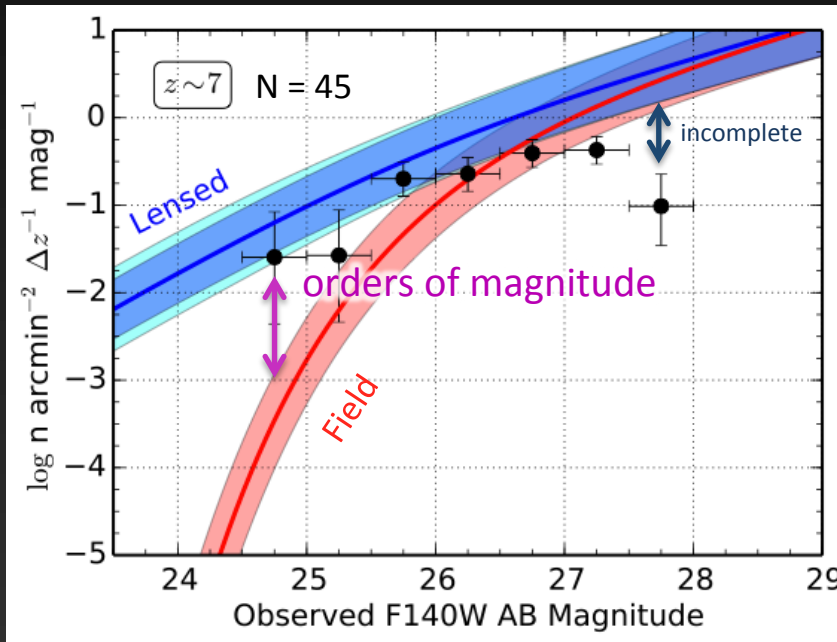
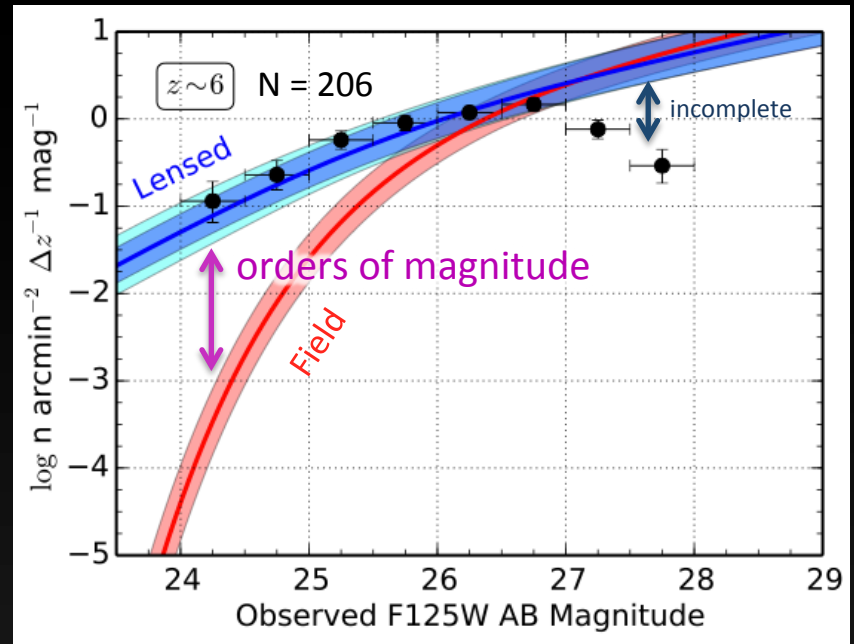
5 selected for lensing strength





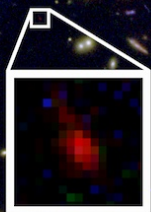
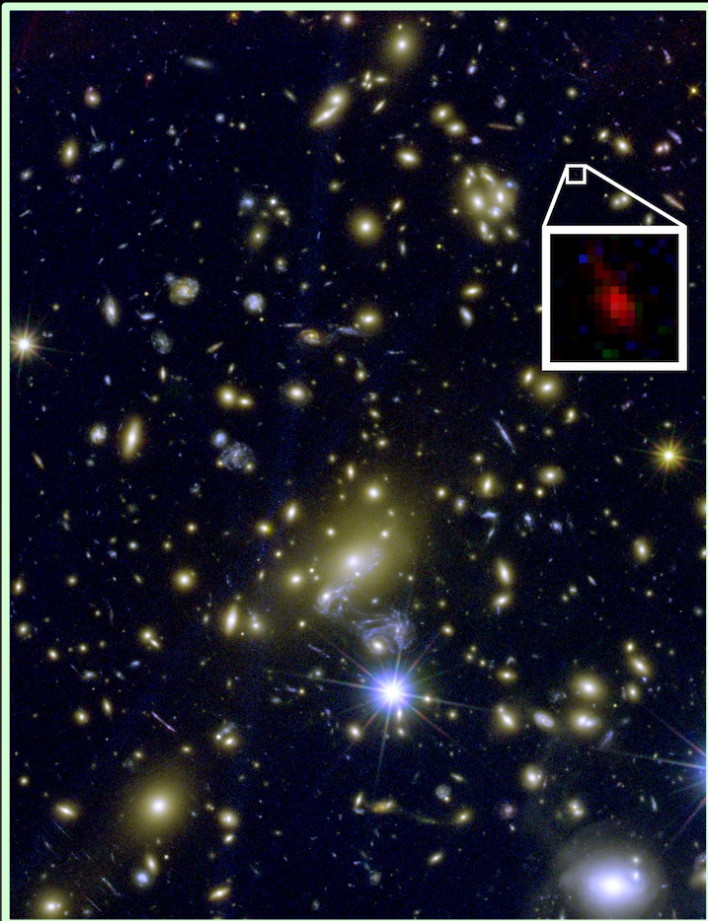
Lensed high-redshift galaxies discovered efficiently

Bradley et al. (2013): 18 / 25 CLASH clusters



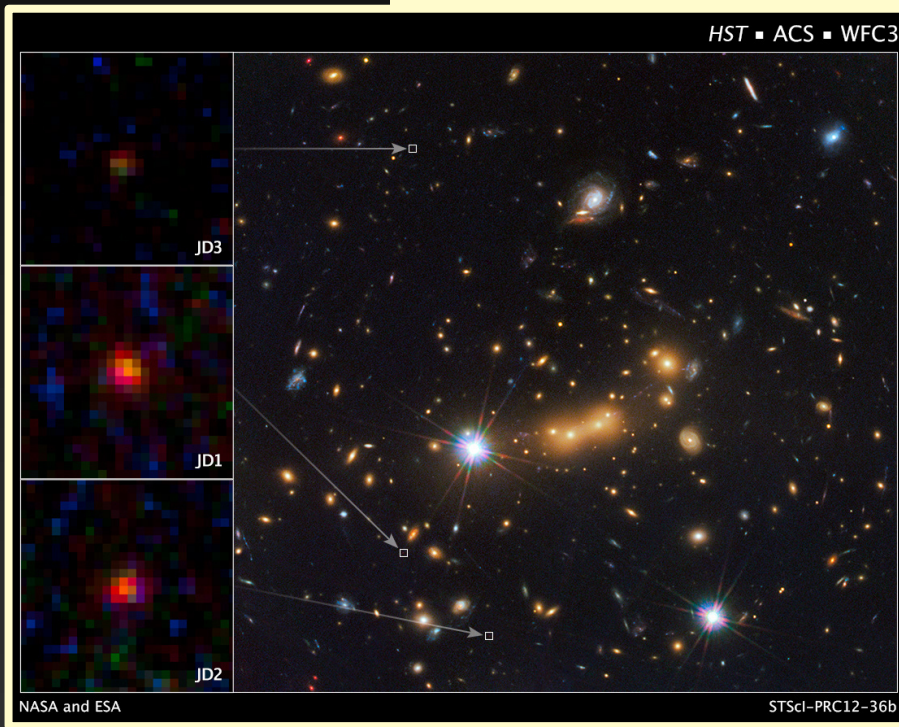


CLASH revealed two robust candidates in the first 500 Myr



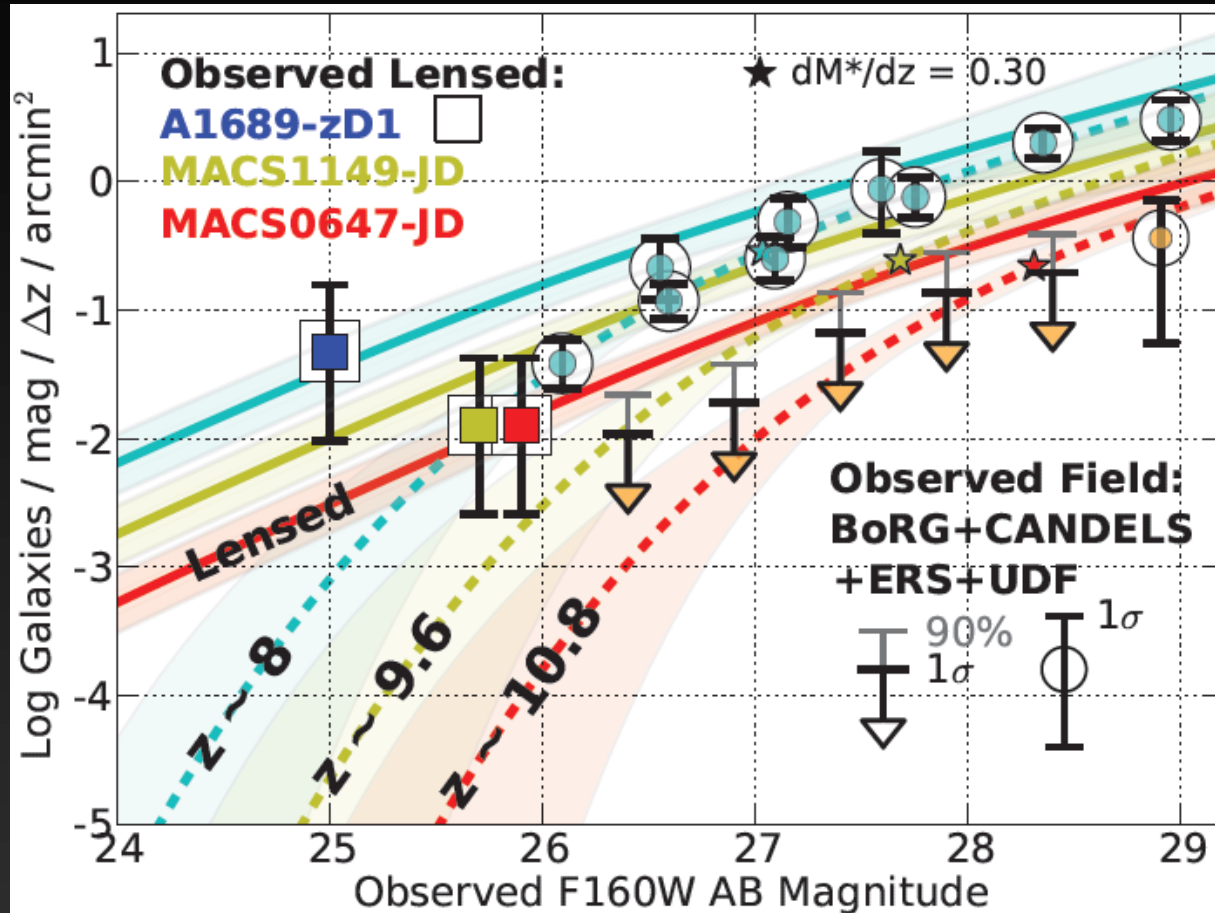
MACS1149-JD
 $z \sim 9.6$ (490 Myr)
Wei Zheng et al. (2012)

MACS0647-JD
 $z \sim 10.8$ (420 Myr)
Dan Coe et al. (2013)



in 2 / 5 "high-magnification" CLASH clusters

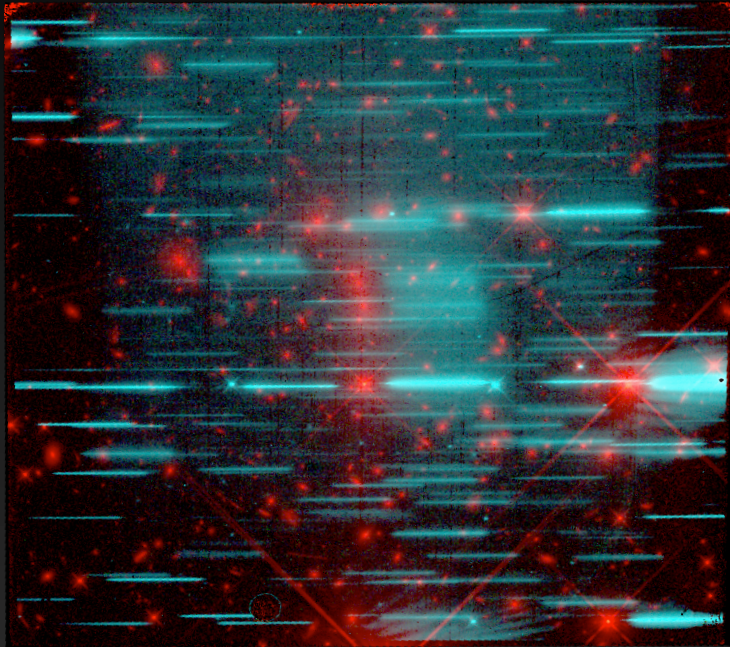
$z \sim 11$ detection consistent with extrapolated lensed LF



Coe+13

See talks by Pirzkal and Lam on $z \sim 11$ HST grism and Spitzer analyses

HST grism 12 orbits (PI Coe)



Spitzer 100 hours (PI Coe)





PI Jennifer Lotz

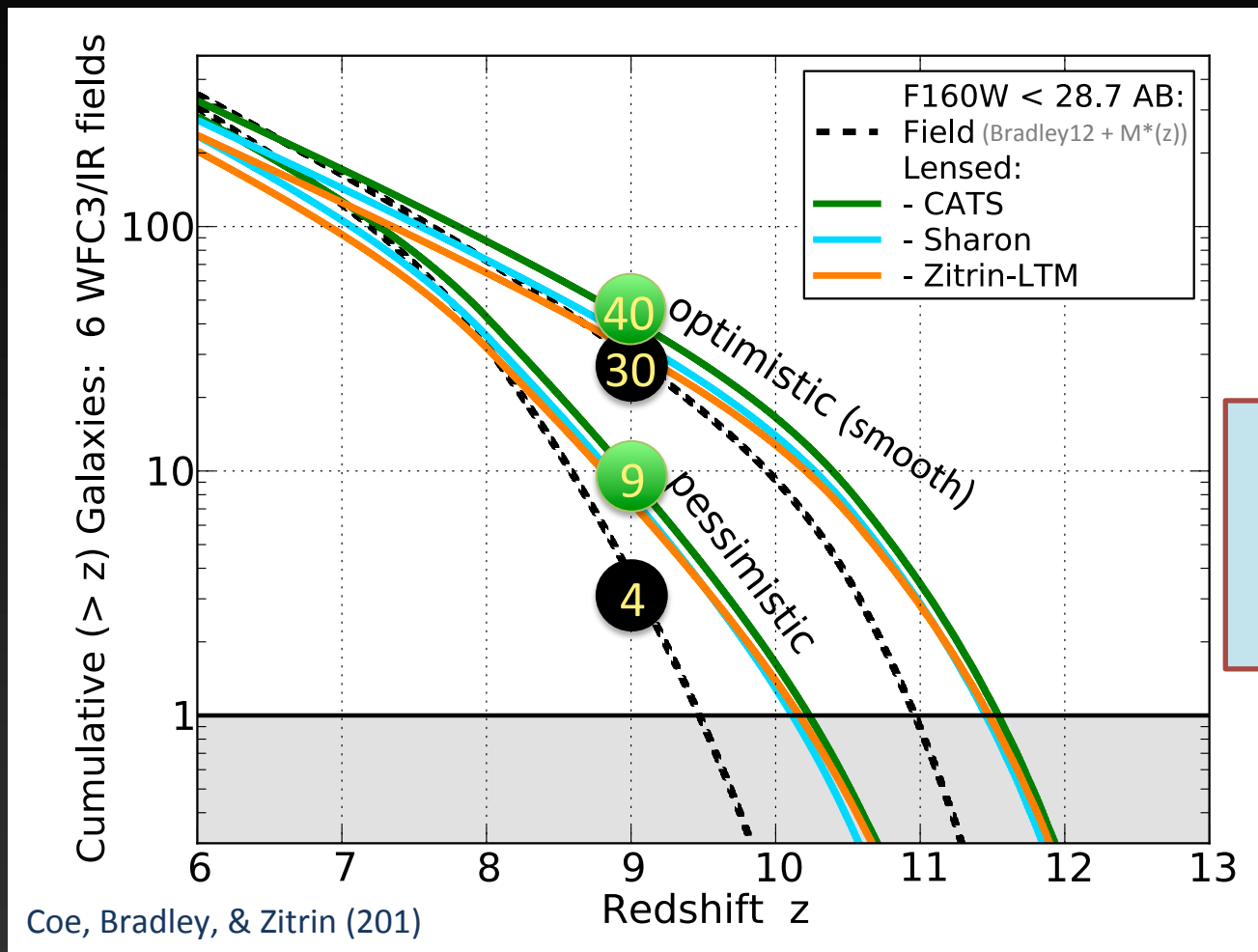


<http://www.stsci.edu/hst/campaigns/frontier-fields/>

Up to 840 Hubble orbits
+ 1000 Spitzer hours
directors' discretionary time



We predict up to ~ 70 $z > 9$ galaxies in the Frontier Fields (6 per field) not accounting for incompleteness



Lens models yield consistent predictions

Frontier Fields lens models available via MAST

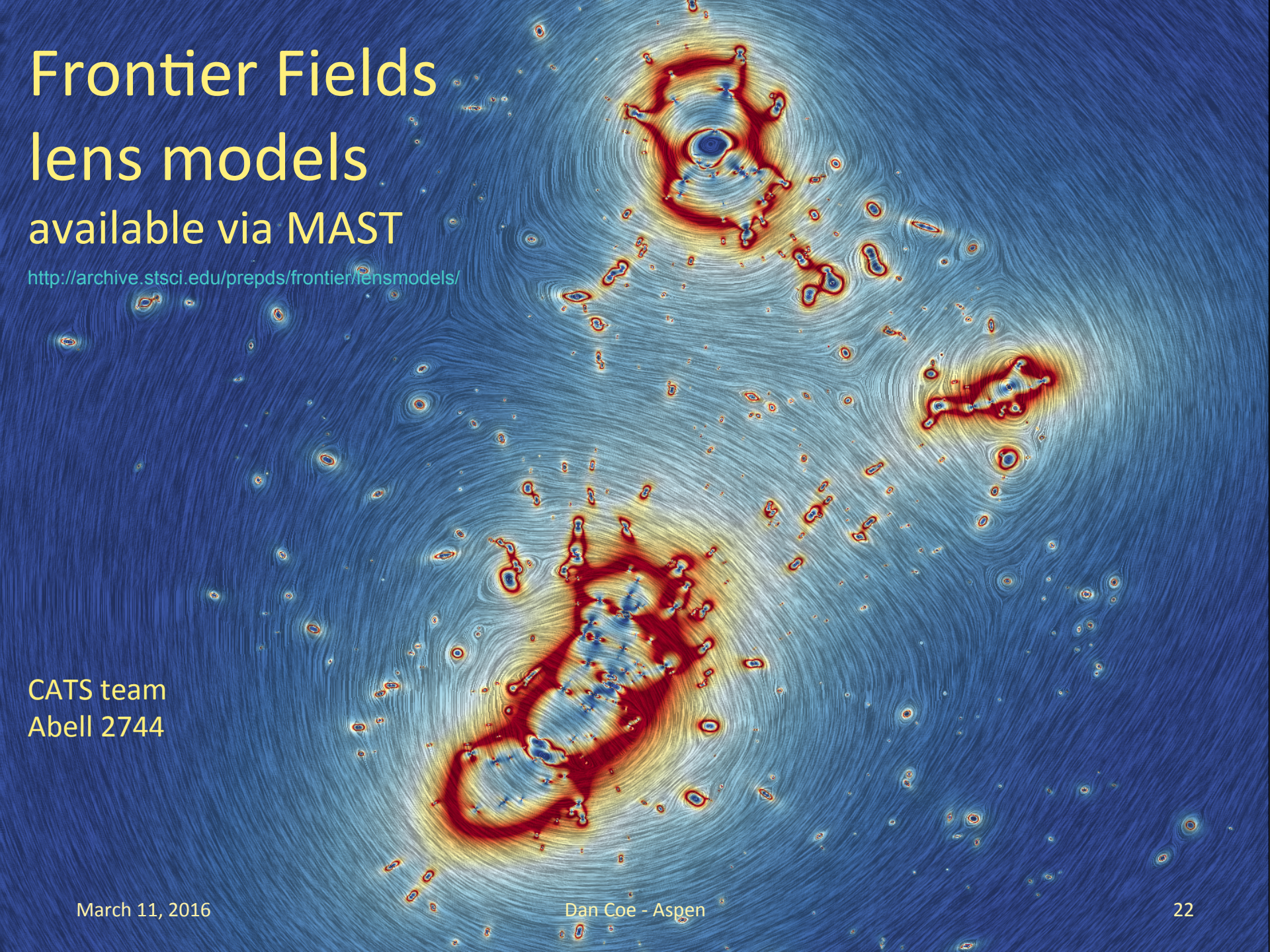
<http://archive.stsci.edu/prepds/frontier/lensmodels/>

CATS team
Abell 2744

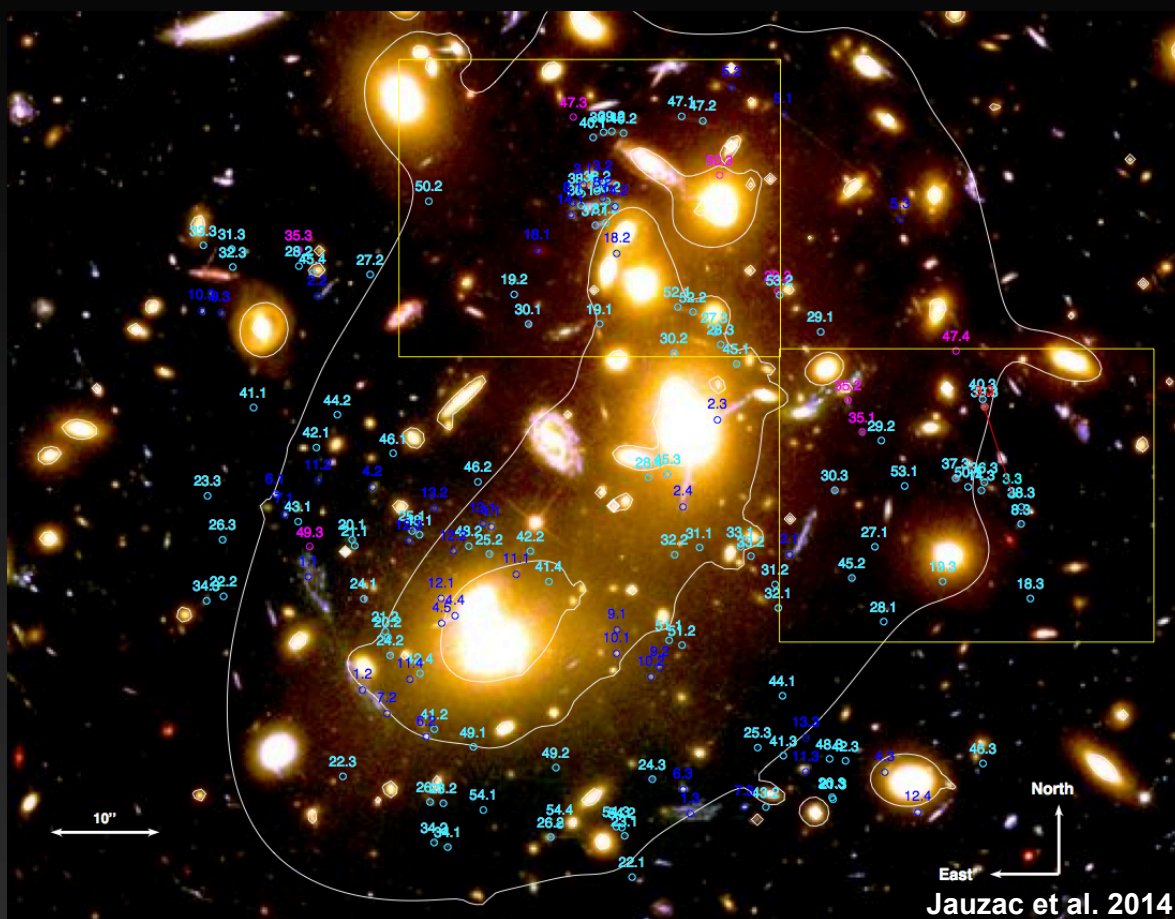
March 11, 2016

Dan Coe - Aspen

22



Best constrained lens models



Abell 2744:

30 pre-FF multiple images

120 new multiple images

7 new but less confident

similar results for other clusters

Interactive model magnification web tool

Used ~2x / day (~1900x)
in first 2+ years
(Nov. 2013 – Mar. 2016)

Hubble Frontier Fields lens model magnification estimates

archive.stsci.edu/prepds/frontier/lensmodels/webtool/magnif.html

Calculated at your input redshift(s) based on the mass and shear maps submitted by each team (see [lensing primer](#)). (Not interpolated / extrapolated from the magnification maps pre-calculated at $z = 1, 2, 4, 9$ available for download.) [Lens model main page](#)

Input
RA, Dec, redshift
of lensed galaxies

Single lensed galaxy:
RA:
Dec:
z =
observed radius (arcseconds):

List of lensed galaxies: RA, Dec, z, (optional) radius

0:14:23.219	-30:23:44.07	10.8
4h16m12.356s	-24h04m35.01s	7.8
(109.36925, 37.72828)	5.3	
11.49 36.888	22.24 18.93	3.2
342.2026	-44.536809	2.1
2:39:51	-1:34:09	0.9
3:47:968,	-30:37:596	9.6

Save results with run number and optional passcode:

(* = Based in part on Frontier Fields imaging. All models are "version 1" unless otherwise noted.)

[Models: \(availability\)](#) % confidence, calculated from a range of models provided by each group
[\(README\)](#) show all results from each range of models, yielding likelihood distributions

- [CATS](#) with uncertainties
- [CATS \(version 2\)*](#) with uncertainties
- [Sharon](#) with uncertainties
- [Sharon \(version 2\)](#) with uncertainties
- [Zitrin-NFW](#) with uncertainties
- [Zitrin-LTM](#) with uncertainties
- [Zitrin-LTM-Gauss](#) with uncertainties
- [GLAFIC*](#) with uncertainties
- [Williams](#) with uncertainties
- [Williams \(version 2\)*](#) with uncertainties
- [Bradac](#) with uncertainties
- [Merten](#) with uncertainties

Uncertainty calculations add a few seconds response time per galaxy per group.

archive.stsci.edu/prepds/f...
archive.stsci.edu/prepds/frontier/lensmodels/webtool/output_000879_JD1C.html

Hubble Frontier Fields lens model magnification estimates

Output saved to [output_000879_JD1C.html](#)

Lensing cluster: Abell 2744 ($z = 0.308$)

Lensed source ($z = 9.8$):
RA, Dec = (00:14:18.607, -30:24:31.36) = (3.57753, -30.40871)
observed radius = 0.0 arcseconds

Output
Magnification estimates
with uncertainties

CATS	2.95 best;	2.99 ^{+0.06} _{-0.05}	[2.93, 3.05]	median and 68.3% confidence range
CATSV2				
Sharon	2.81 best;	2.70 ^{+0.20} _{-0.11}	[2.60, 2.90]	median and 68.3% confidence range
Sharonv2	2.92 best;	2.86 ^{+0.16} _{-0.08}	[2.78, 3.02]	median and 68.3% confidence range
Zitrin-NFW	2.95 best;	3.65 ^{+1.13} _{-0.76}	[2.89, 4.78]	median and 68.3% confidence range
Zitrin-LTM	37.07 best;	5.41 ^{+4.60} _{-1.81}	[3.60, 10.01]	median and 68.3% confidence range
Zitrin-LTM-Gauss				
GLAFIC	3.08 best;	3.02 ^{+0.17} _{-0.18}	[2.84, 3.18]	median and 68.3% confidence range
Williams	1.48 best;	1.52 ^{+0.17} _{-0.19}	[1.33, 1.69]	median and 68.3% confidence range
Williamsv2				
Bradac	4.59 best;	3.45 ^{+0.37} _{-0.41}	[3.04, 3.82]	median and 68.3% confidence range
Merten	2.76 best;	7.76 ^{+14.29} _{-3.03}	[4.74, 22.05]	median and 68.3% confidence range

For a given (RA, Dec, z),
outputs magnification
estimates and uncertainties
from all models

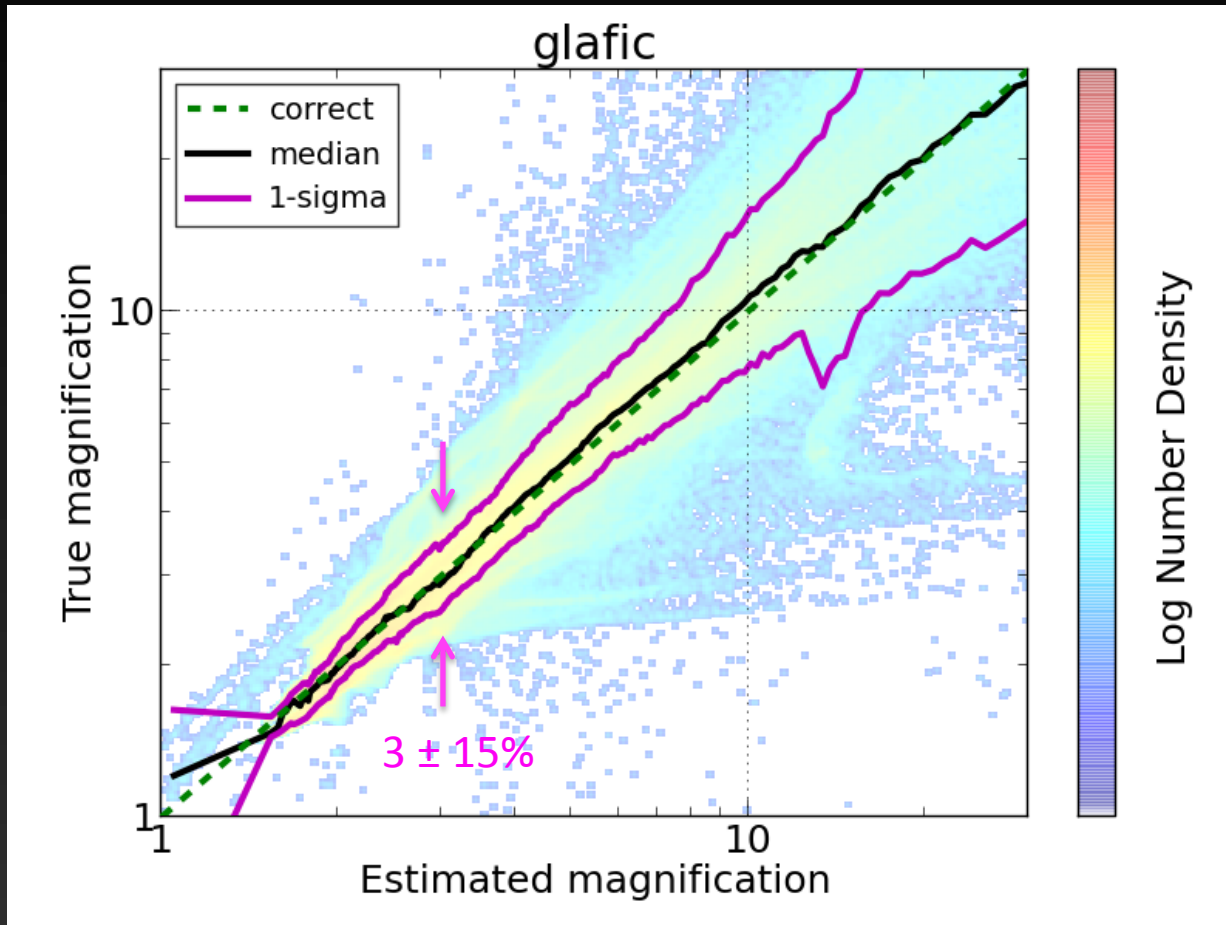
Simulated cluster lensing challenge to quantify model accuracies

**“Hera” simulated
by Meneghetti**

Project led by:
Priyamvada Natarjan
Massimo Meneghetti
Dan Coe

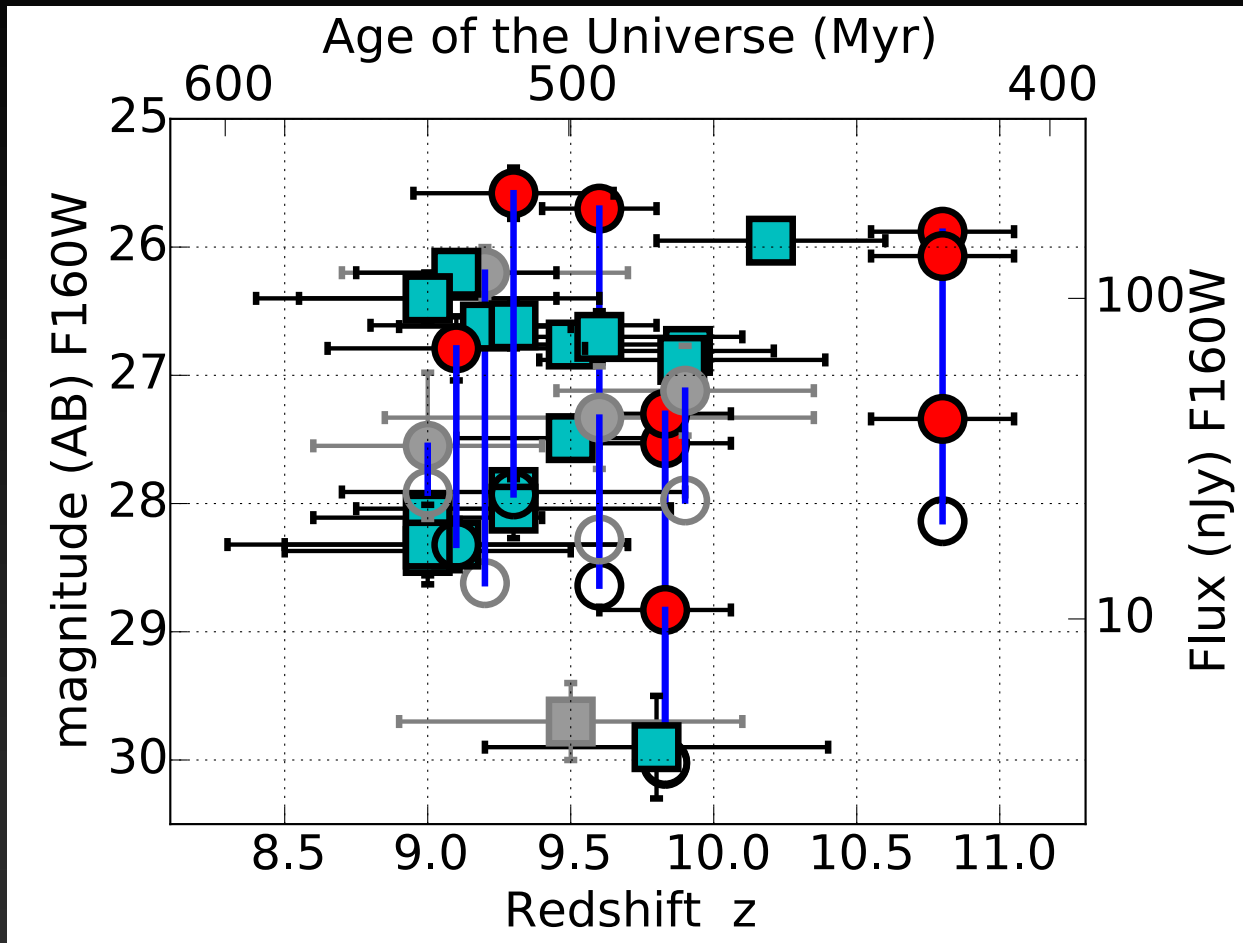


Magnification accuracies quantified; More tests required



“Only” 32 $z > 9$ candidates

vs.
 ~200 $z \sim 8$
 candidates
 (e.g.,
 Bouwens14,
 Finkelstein14)



Lensed
 magnifications

Robust
 Less secure

(Zheng12, Coe13,
 Bouwens13, Zitrin14,
 McLure16)

UDF CANDELS

Field

Robust
 Less secure

(Ellis13, Oesch13,
 Oesch14, Bouwens15,
 McLure16)

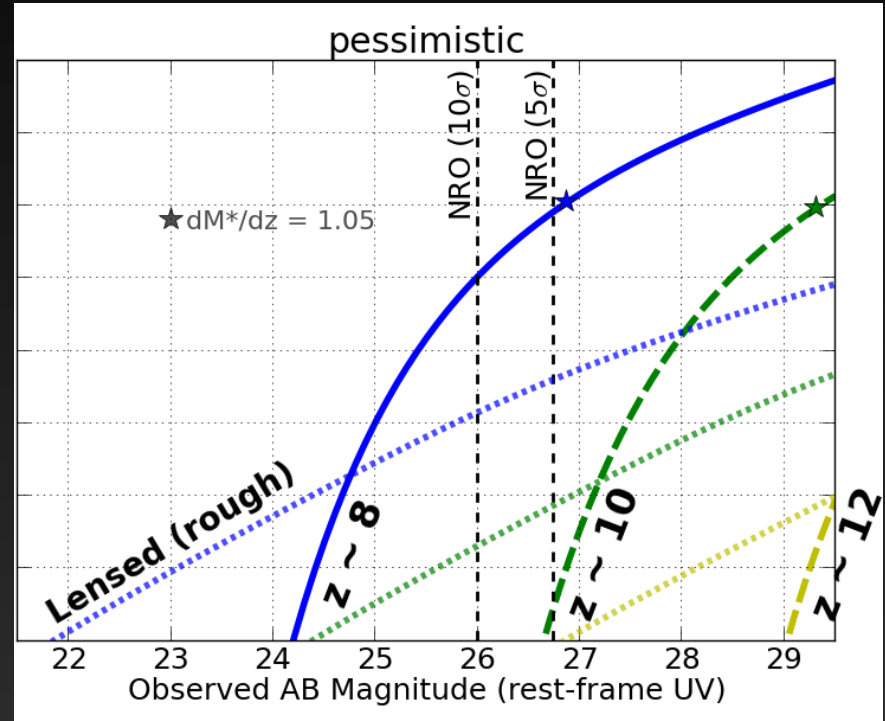
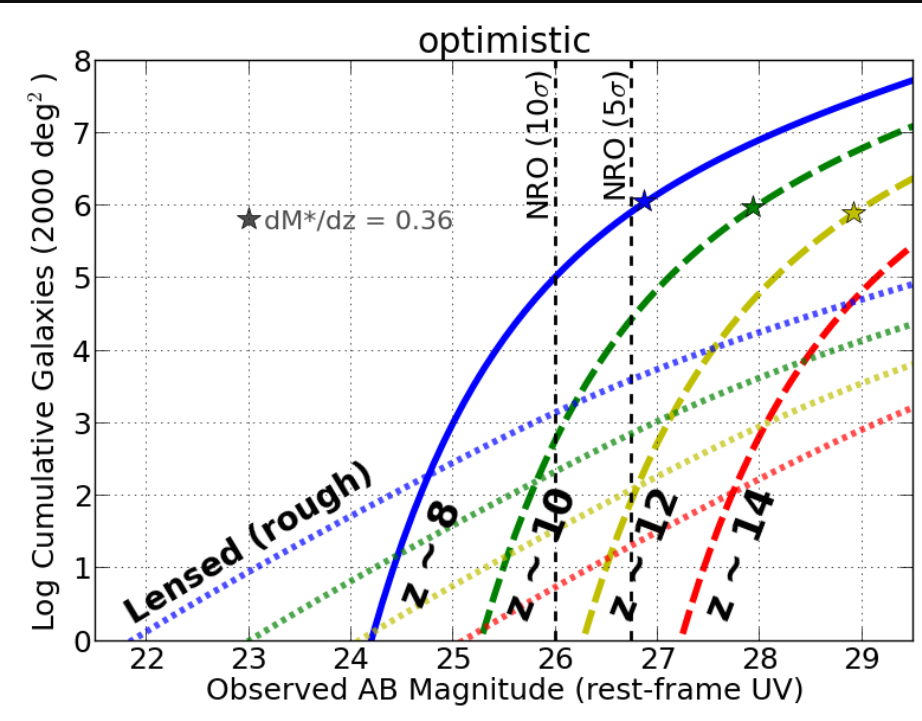
Beyond the Frontiers...

Euclid and WFIRST will yield many high-z candidates

Would be most useful before or during JWST's mission

WFIRST: 100,000's at $z \sim 8$ and perhaps some out to $z \sim 15$

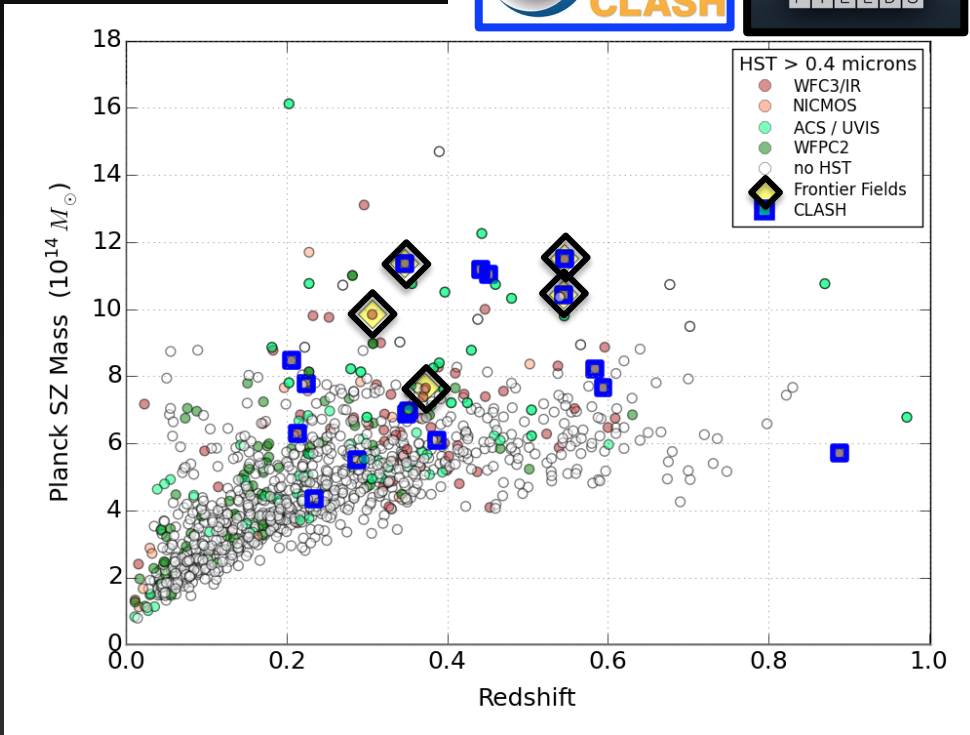
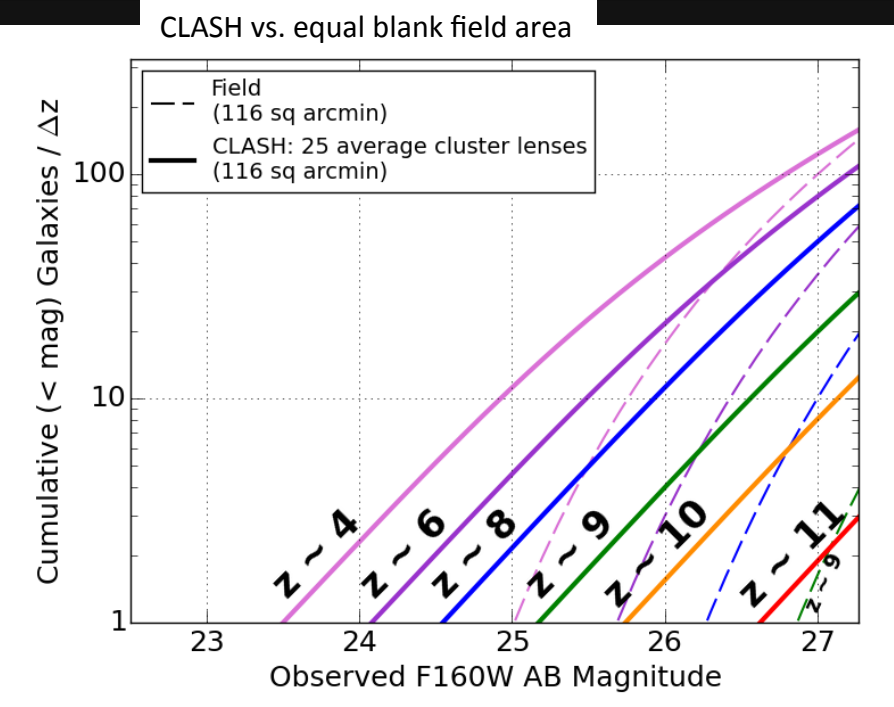
roughly assuming one strong lensing cluster per square degree



D. Coe & L. Bradley
WFIRST white paper submission

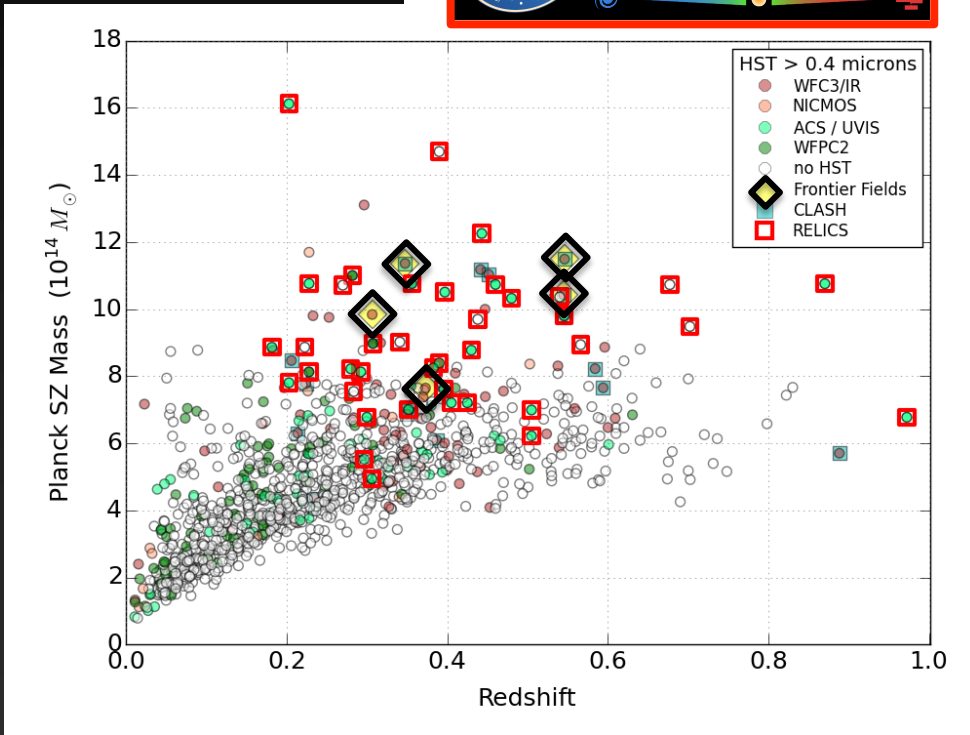
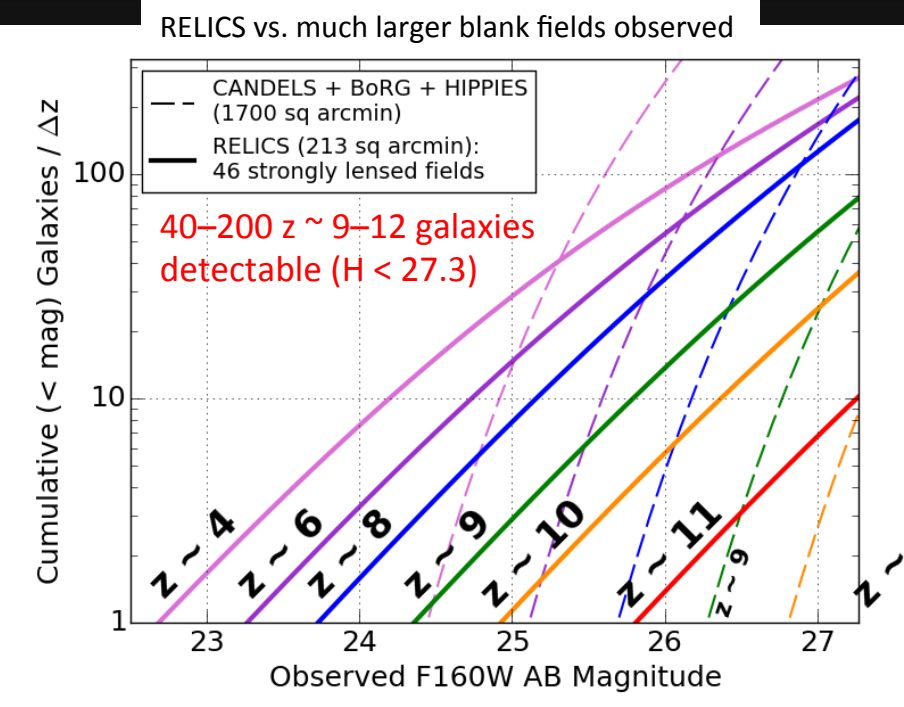
Beyond the Frontiers...

Hubble can discover the best and brightest high-redshift candidates in time for JWST with a large new cluster survey



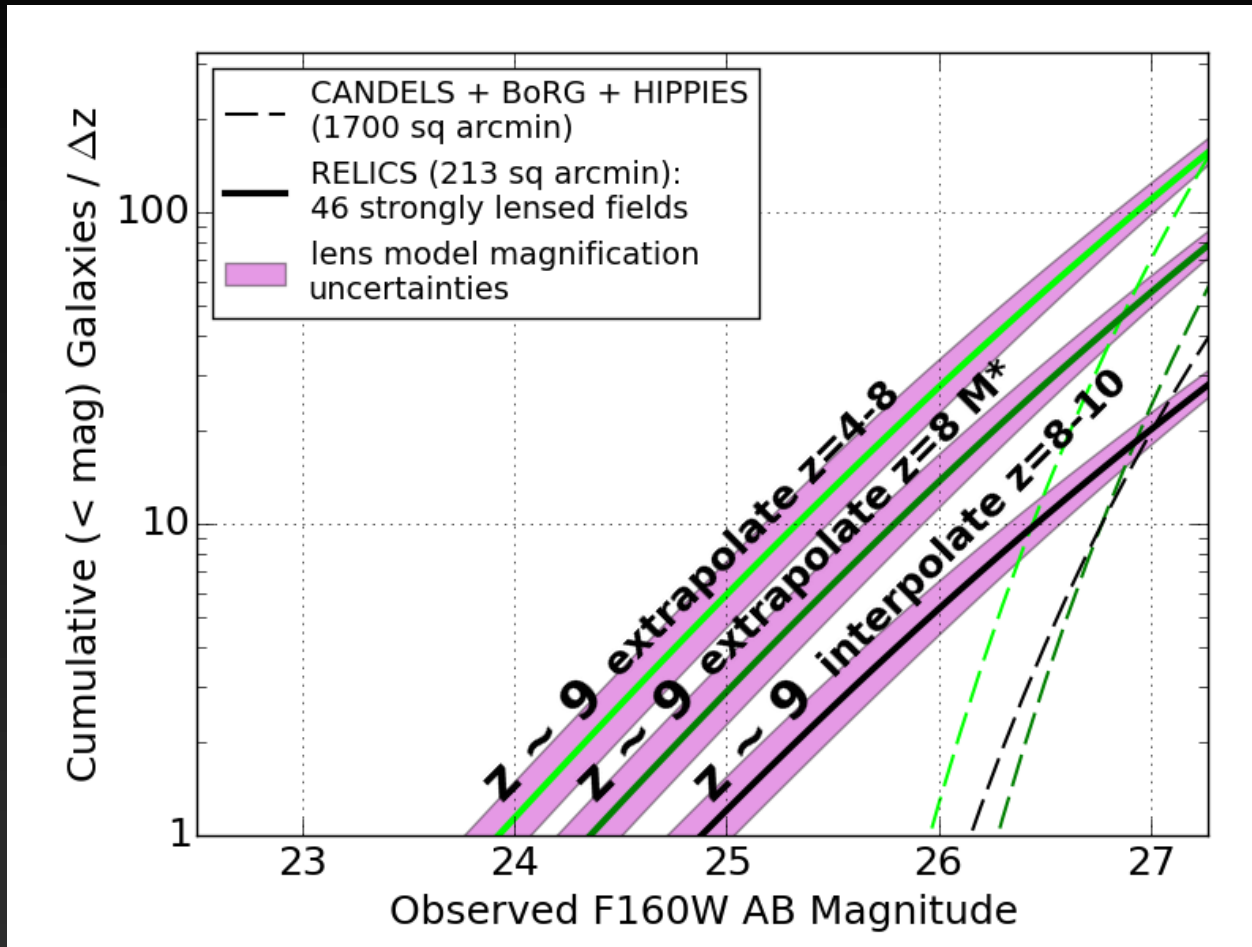
Beyond the Frontiers...

Hubble can discover the best and brightest high-redshift candidates in time for JWST with a large new cluster survey



$z \sim 9$ efficiency and constraints

$z \sim 9 - 12$ galaxies
if complete to $H < 27.3$



200
100
40

RELICS: Reionization Lensing Cluster Survey

Observations

46 fields lensed by 41 clusters
HST: 190 orbits + 77 parallel (incl. 20 for SN)
- 3 orbits ACS (minus archival)
- 2 orbits WFC3/IR Frontier Fields filters
Spitzer: 390 hours (PIs Bradac, Soifer)

Science

high-redshift galaxies
cluster mass scaling relations
merger physics + DM constraints
supernovae

Delivery

no proprietary period HST images
reduced images + catalogs
2-3 months after completion of each field
final high-z candidates + lens models
Nov. 2017 (JWST GO call for proposals)

Example

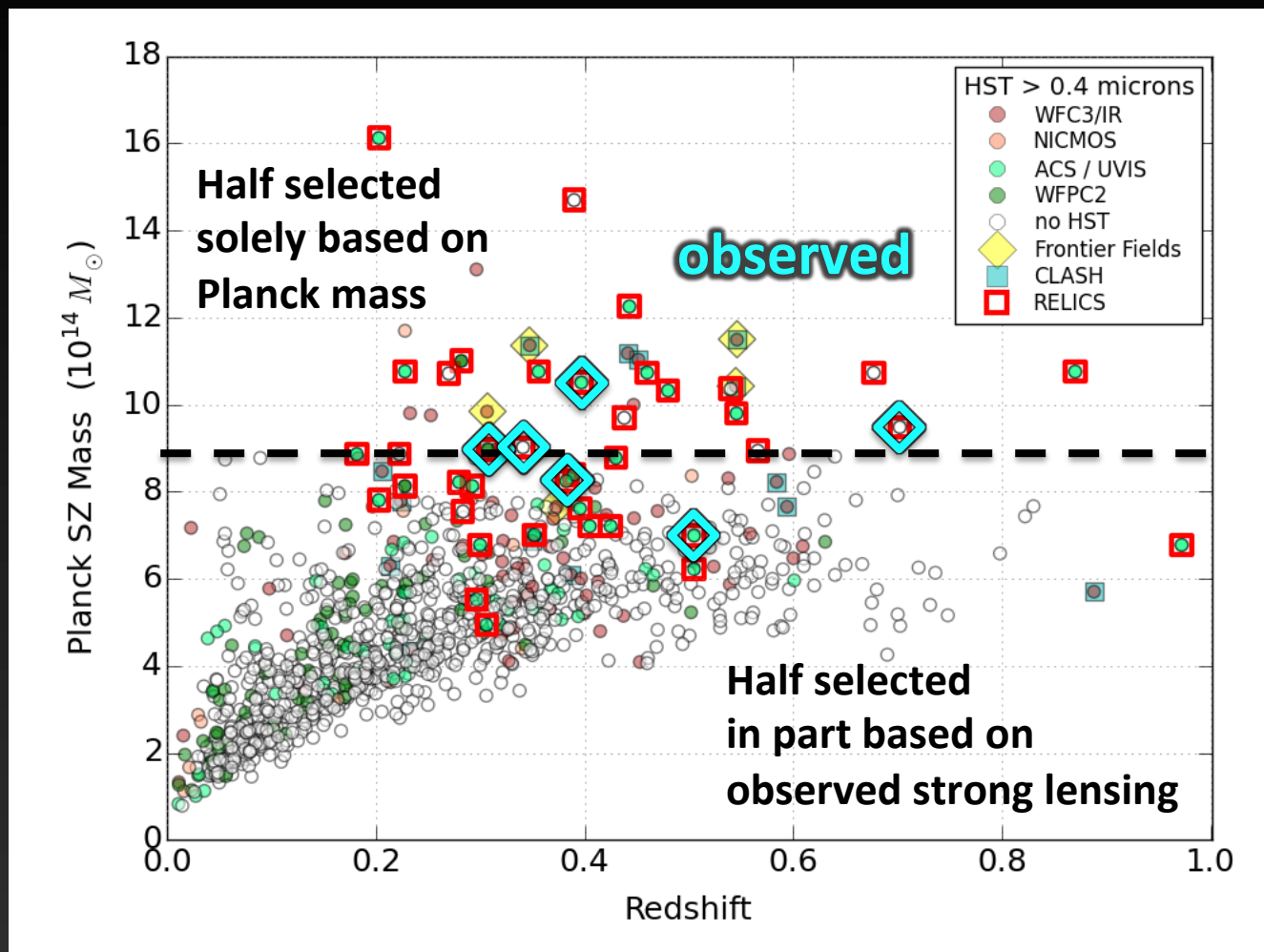
ACS imaging of A2163, the most massive cluster according to Planck



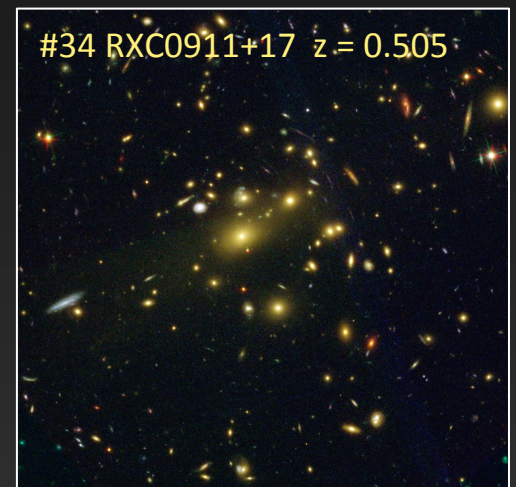
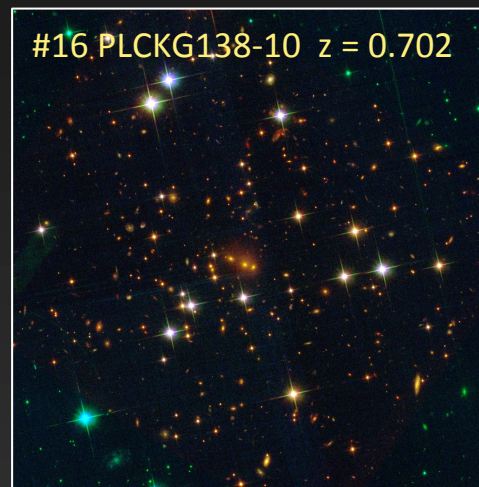
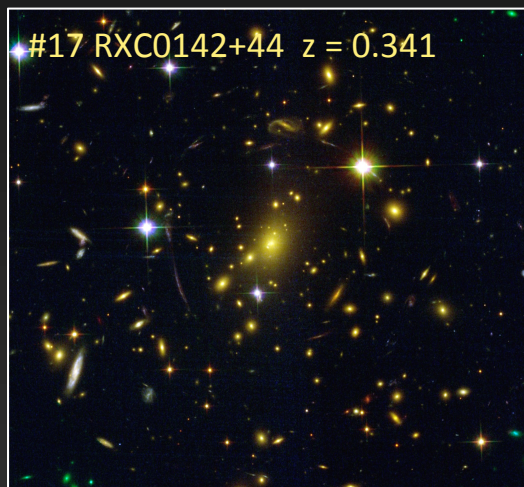
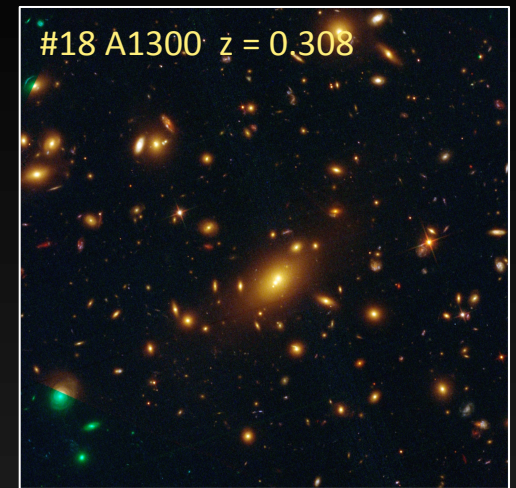
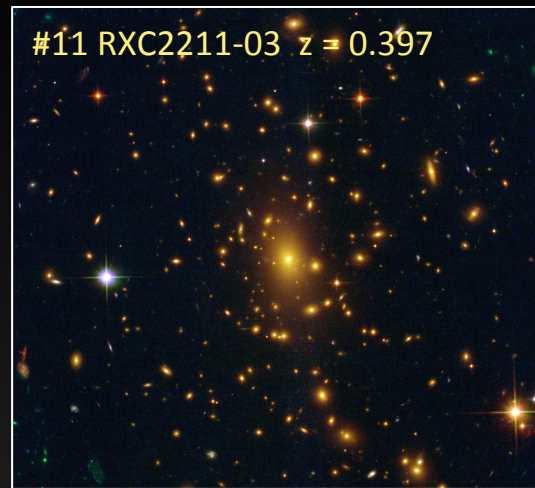
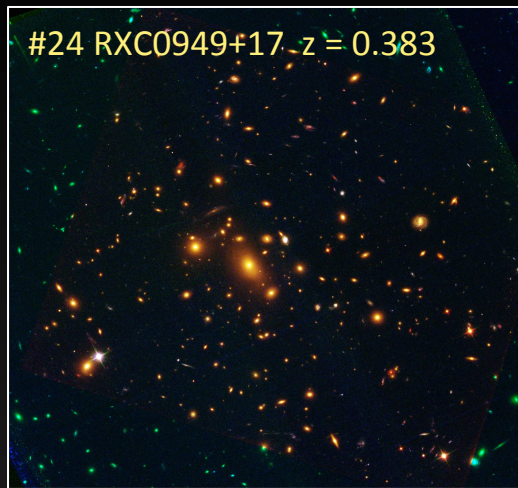
RELICS will obtain the first WFC3/IR imaging

Dan Coe (PI)
Larry Bradley (Deputy PI)
Felipe Andrade-Santos
Roberto Avila
Rychard Bouwens
Maruša Bradač
Daniela Carrasco
Nicole Czakon
Will Dawson
Brenda Frye
Austin Hoag
Kuang-Han Huang
Traci Johnson
Christine Jones
Daniel Lam
Ramesh Mainali
Cordell Newmiller
Pascal Oesch
Sara Ogaz
Rachel Paterno-Mahler
Adam Riess
Steve Rodney
Russell Ryan
Brett Salmon
Irene Sendra-Server
Keren Sharon
Dan Stark
Lou Strolger
Michele Trenti
Keiichi Umetsu
Benedetta Vulcani
Adi Zitrin

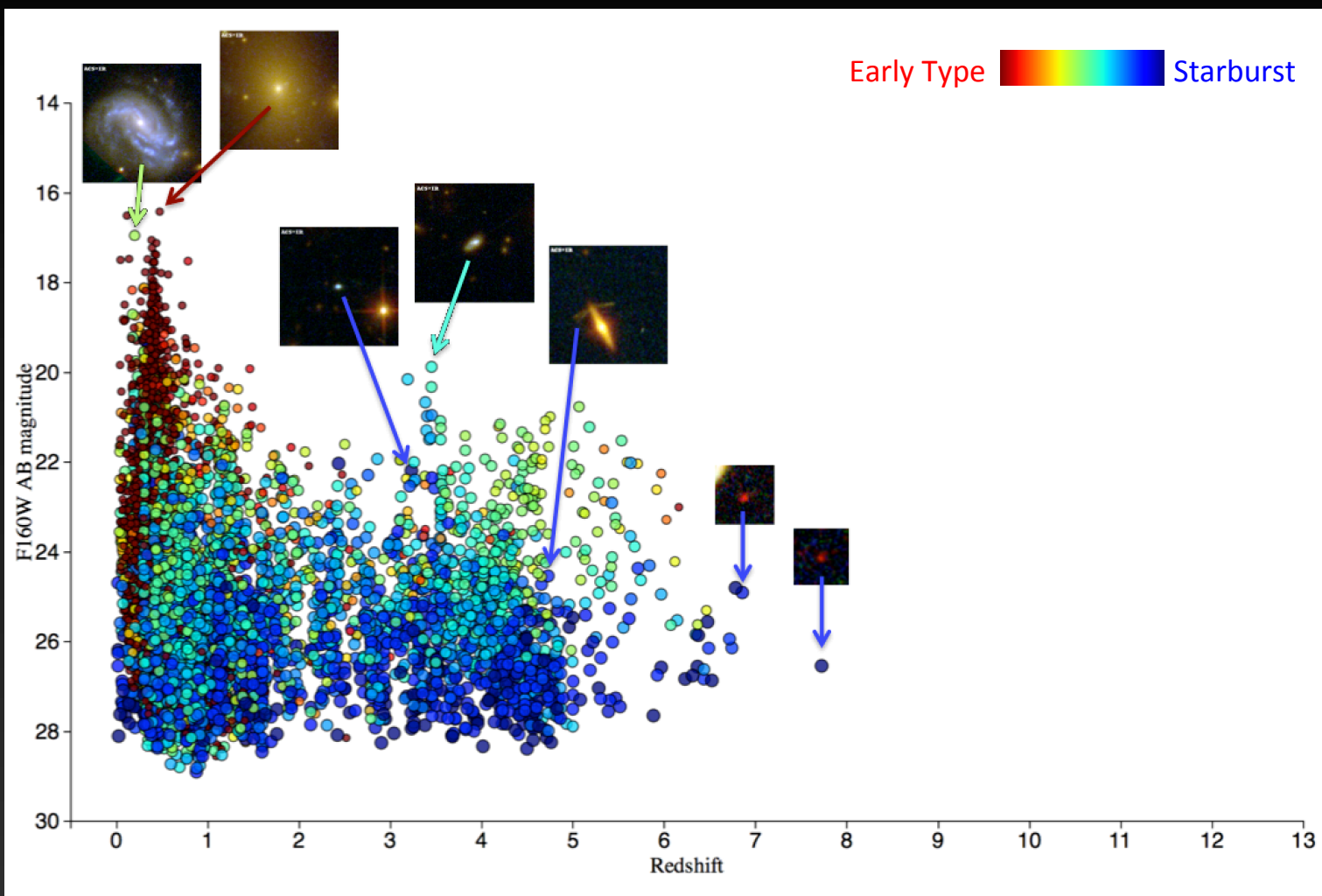
High Planck mass best for lensing?



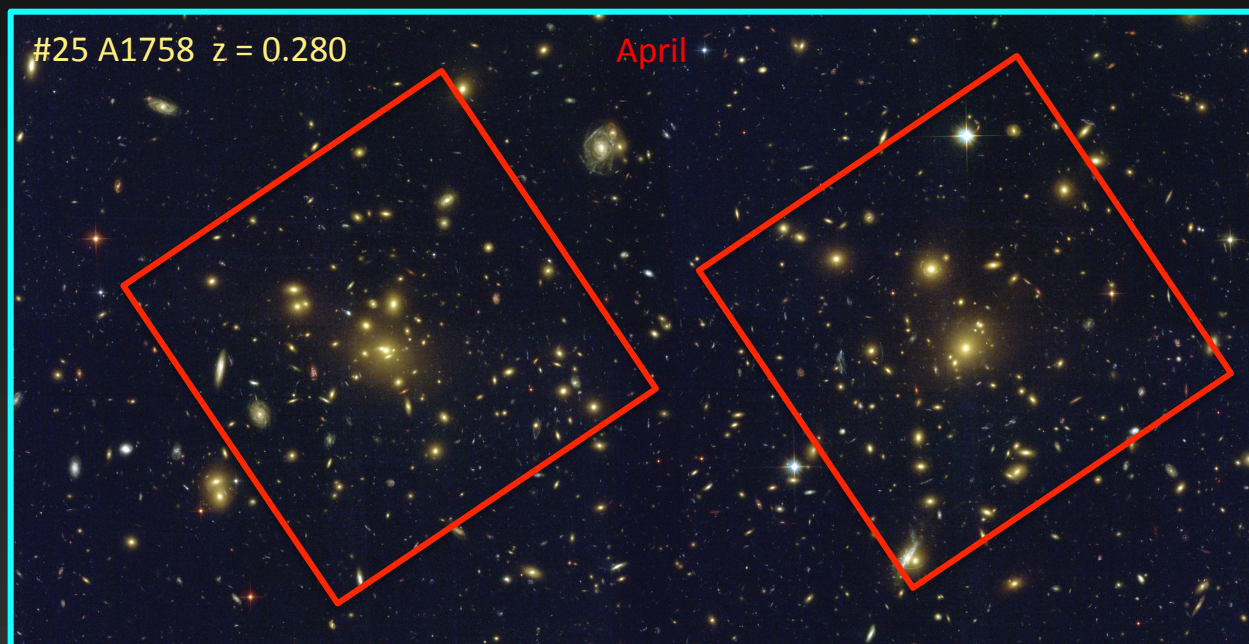
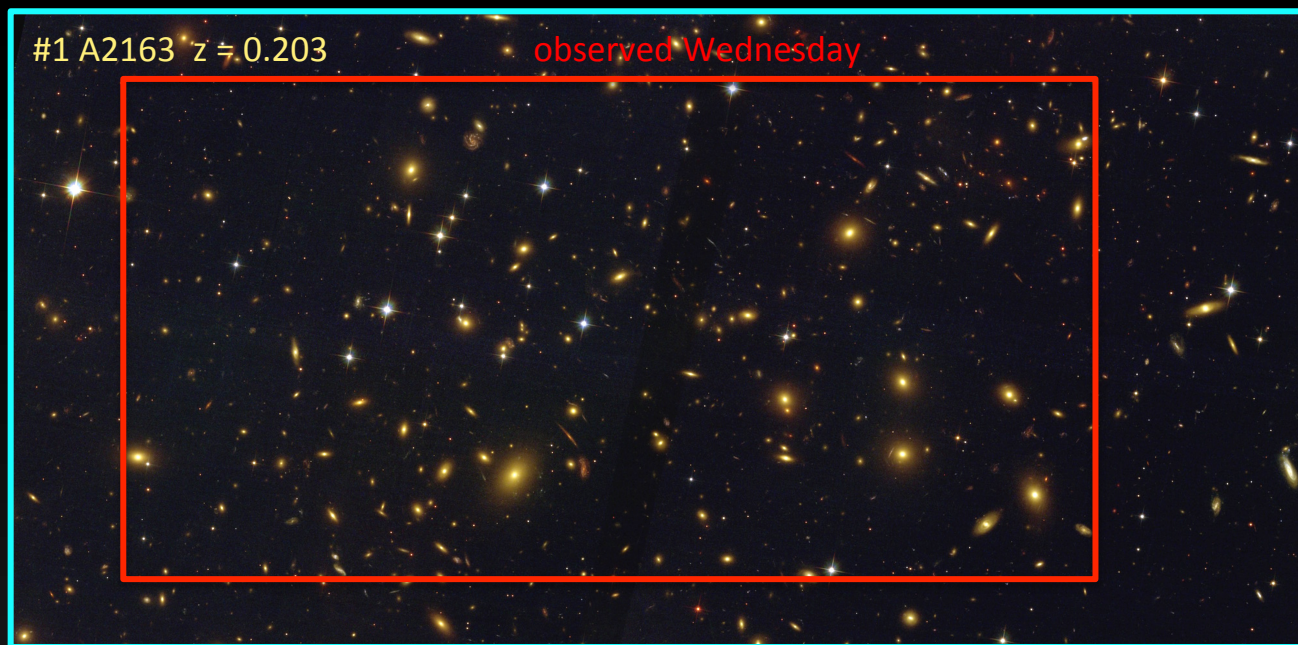
6 / 41 RELICS clusters observed so far



Galaxies in first 6 RELICS cluster images



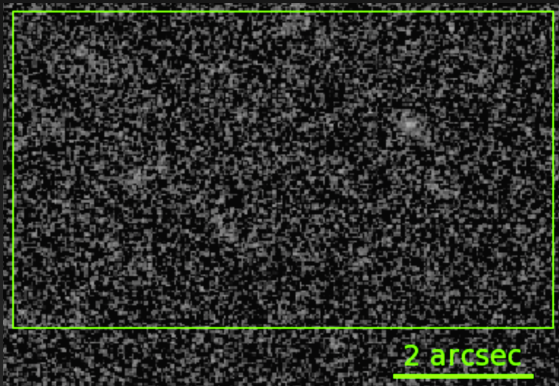
The best
RELICS
clusters
are yet
to come



Maximum Likelihood Detection

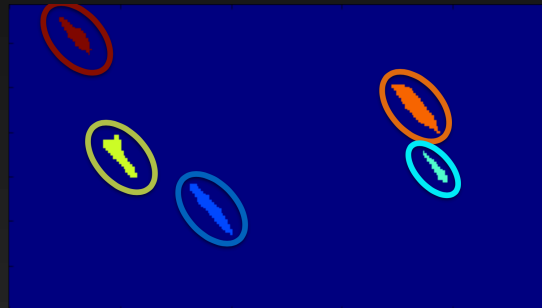
Will Dawson with Jim Bosch and Michael Schneider

2-orbit image

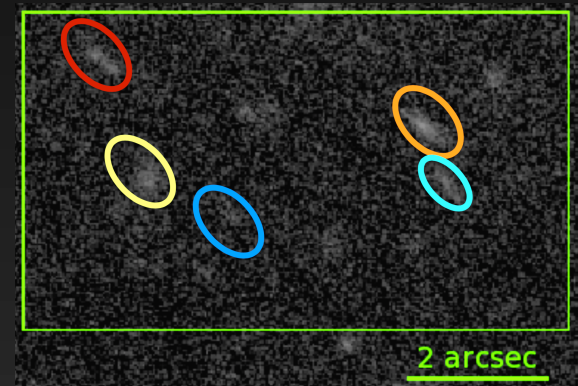


MACS1149: CLASH

Detection
using sheared kernel



15-orbit image



MACS1149: Frontier Fields

Oct 2015
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RELICS
Oct 2015
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Apr 2017



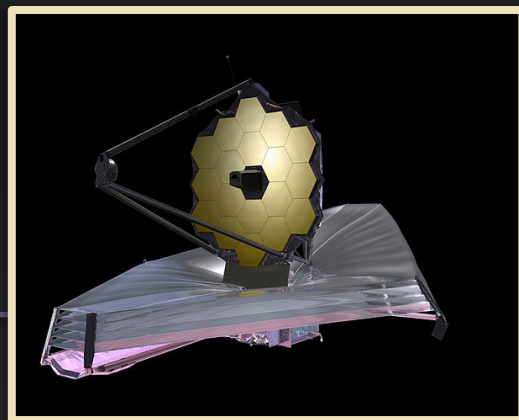
Find the best and brightest high-redshift candidates now in time for JWST

GTO
Jan 2017 call
Apr 2017 deadline
Jun 2017 targets public

ERS (public data)
Mar 2017 call
Jul 2017 deadline
Sep 2017 TAC

(some proposal dates subject to change)

GO Cycle 1
Nov 2017 call
Feb 2018 deadline
May 2018 TAC



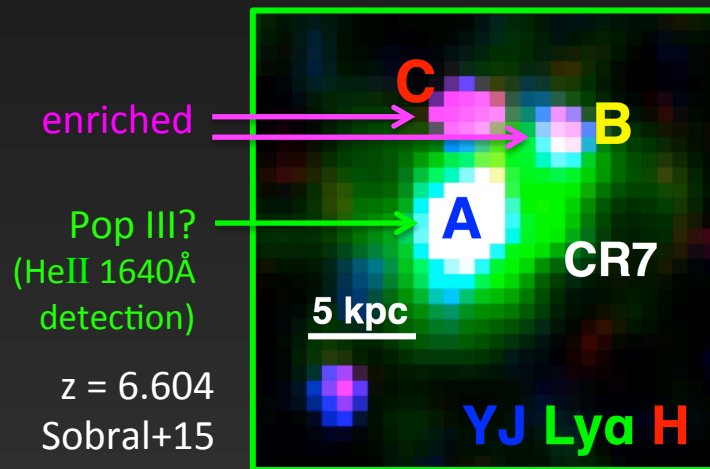
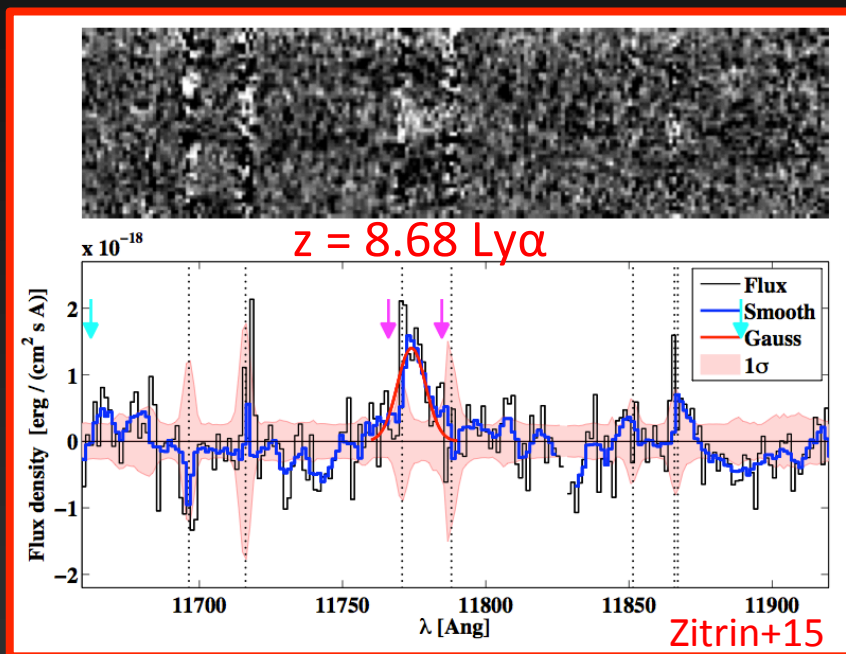
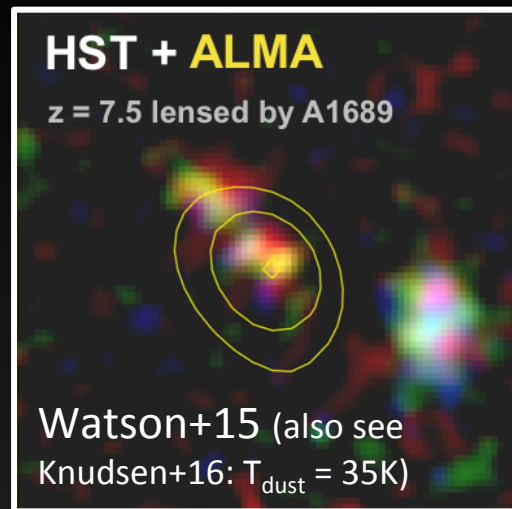
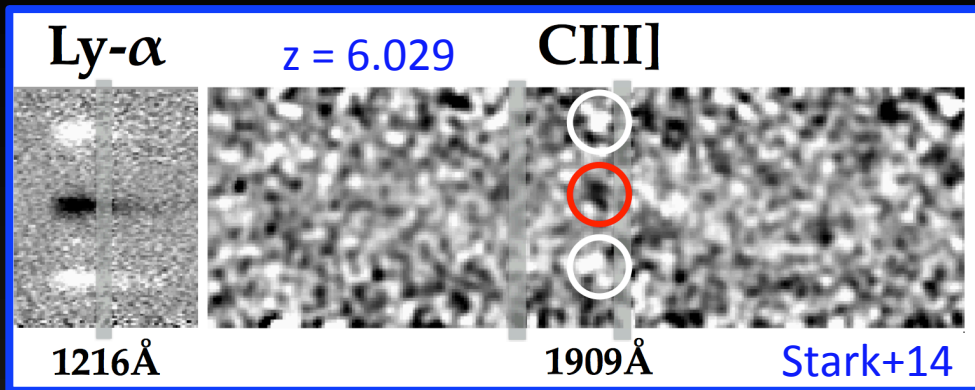
Launch! Oct 2018
commissioning 6 months

Science! Apr 2019
5 – 10 years

March 11, 2016

Dan Coe - Aspen

Confirmations and properties of brightly observed high-z galaxies

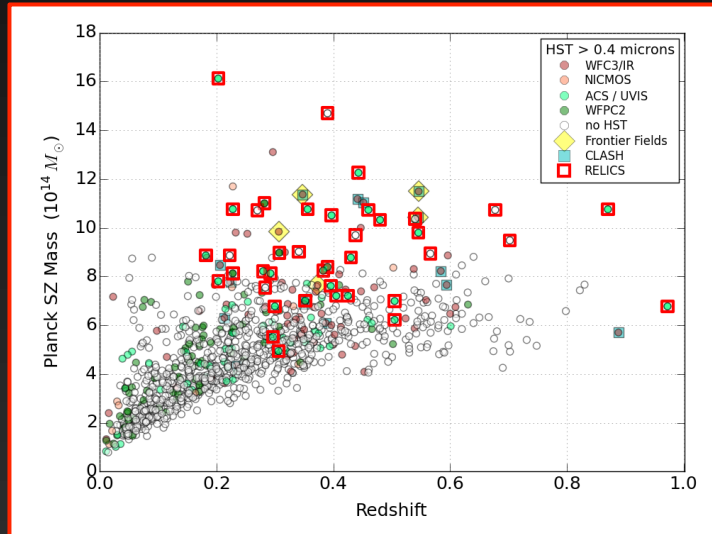
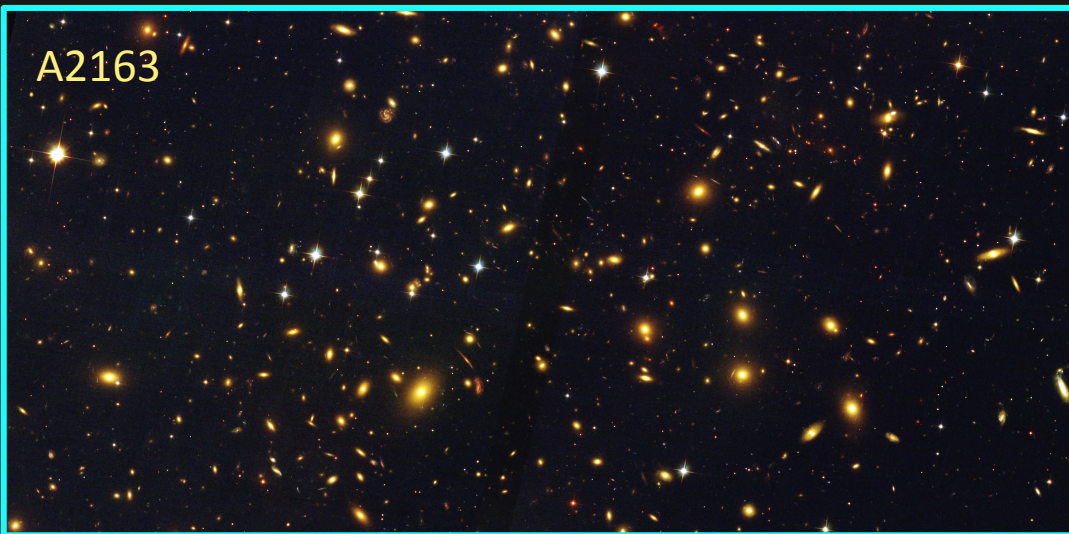


RELICS will deliver more of the “best and brightest” high-redshift candidates in time for JWST

Lensing delivers more, intrinsically fainter, and higher resolution $z > 9$ galaxies



<http://relics.stsci.edu>



The history of astronomy is a history of receding horizons.
— Edwin Hubble

ZwCL 0947+1723

slide credit:
Larry Bradley

