Quantitative Morphological Classifications from z = 0 to z = 3 with SDSS, ACS, and WFC3

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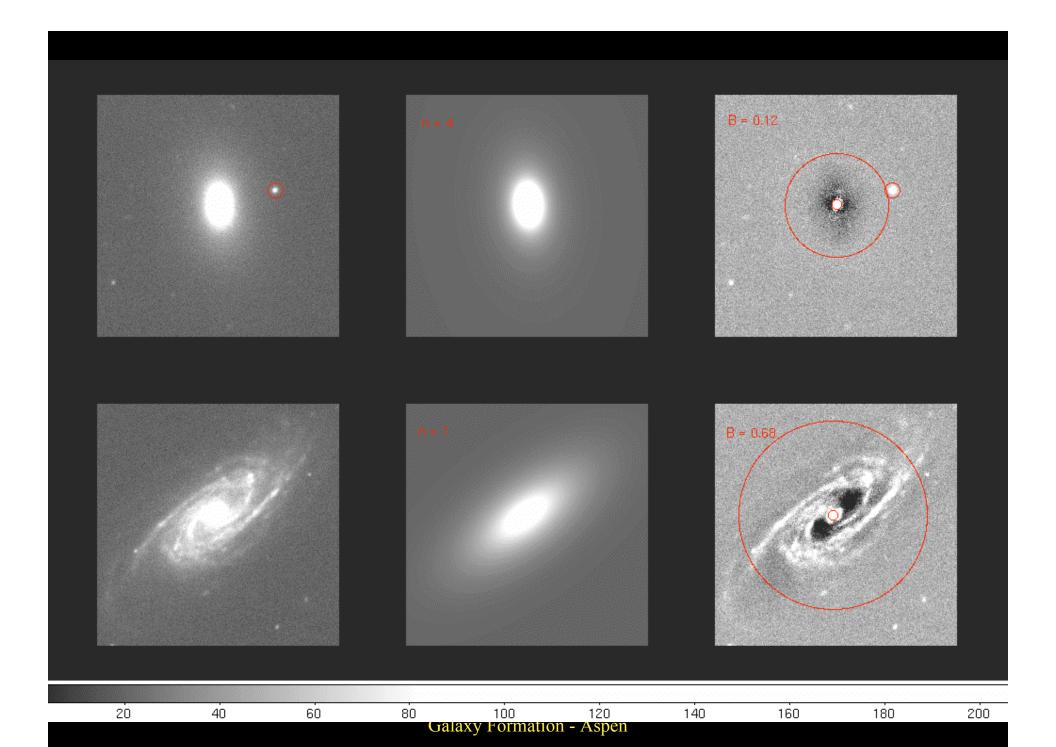
Brad Holden; Marijn Franx; Garth Illingworth; Holland Ford; John Blakeslee; Marc Postman; Dan Kelson; Andrew Zirm; Ivo Labbé and the ACS team

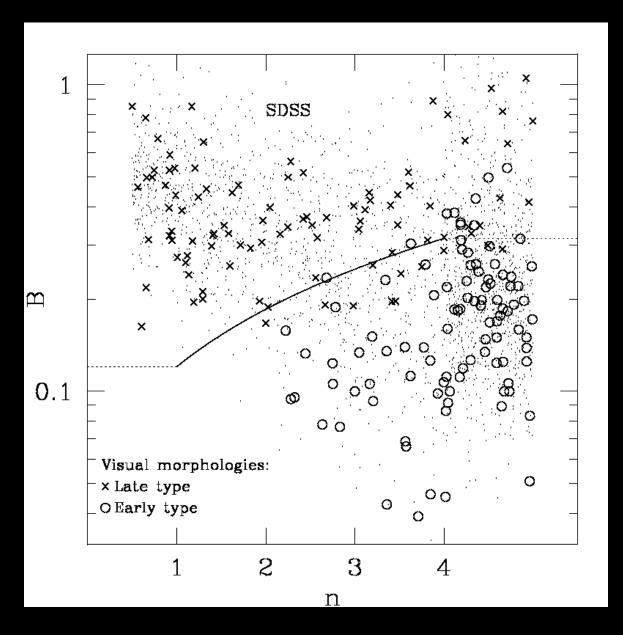
Outline

- Morphologies in the SDSS
- Evolution of the morphology density relation
- What WFC3 will bring

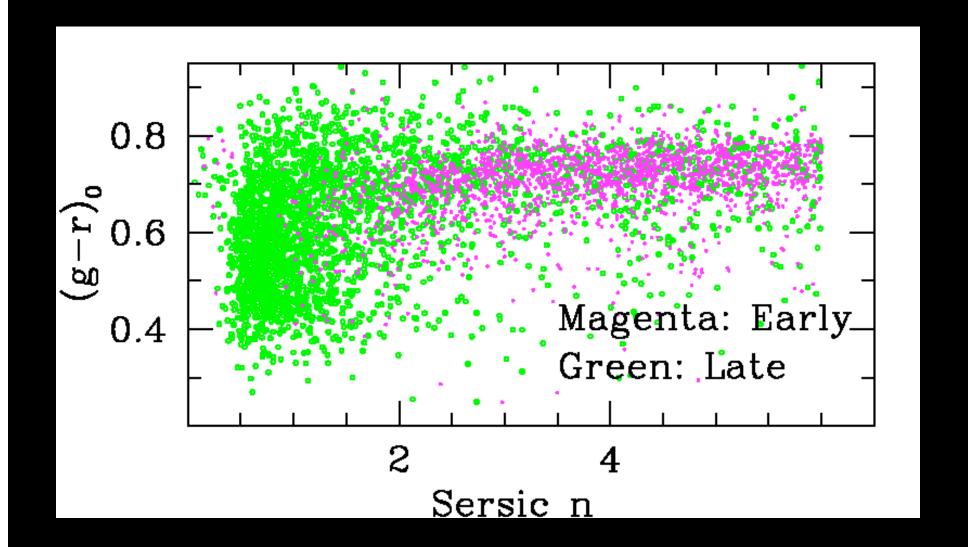
Sample extracted from SDSS (DR5)

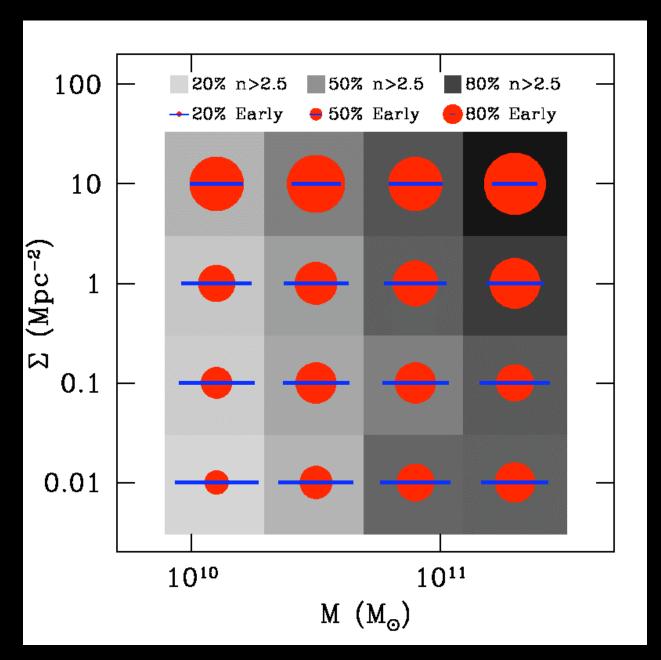
- 0.02 < z < 0.03
- $M_* > 10^{10} \text{ Msol (Bell et al. 2003)}$
- Environment (7th nearest neighbor)
- Morphology
 - Sersic profile fits to g-band images
 (GALFIT; Peng 2002)
 - 'Bumpiness' (Blakeslee et al. 2006)





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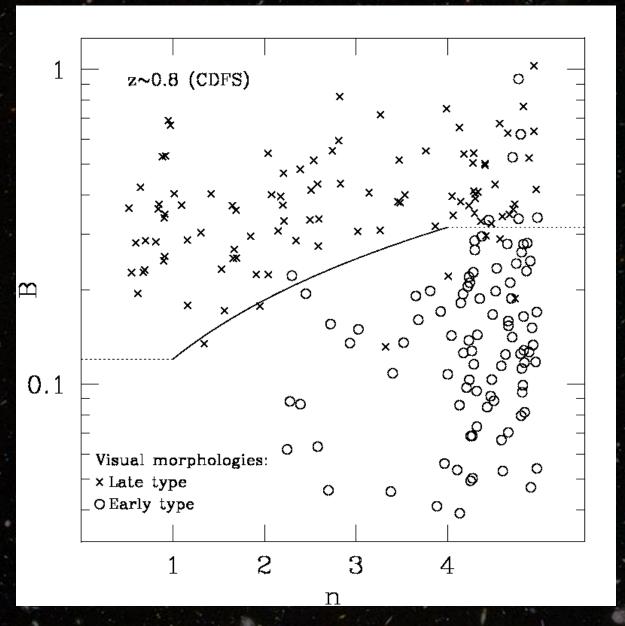
Conclusion

van der Wel 2008 (ApJL, in press, arXiv:0801.1995)

'Structure' and 'morphology' are intrinsically different quantities that behave differently as a function of mass and environment.

Field sample

- SDSS at 0.020 < z < 0.045: 2003 galaxies with M > 4 × 10¹⁰ Msol
- GOODS-South (CDFS) at 0.6 < z < 1.0: 207 galaxies with M > 4 ×10¹⁰ Msol

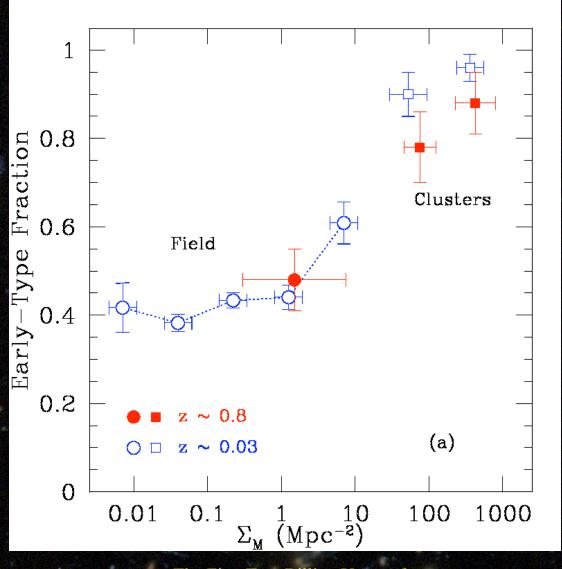


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Cluster sample

- Coma cluster at z = 0.023: 95 galaxies with $M > 4 \times 10^{10}$ Msol
- MS1054 & CL0152 at z = 0.83: 191 galaxies with $M > 4 \times 10^{10}$ Msol

The MDR at z = 0 and z = 0.8



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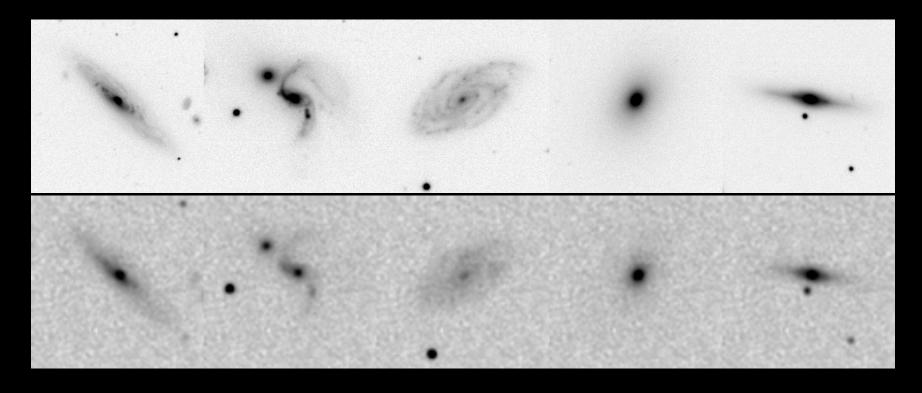
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Conclusion

Holden et al. 2007 (ApJ 670, 190) van der Wel et al. 2007 (ApJ 670, 206)

The early-type galaxy fraction does not significantly evolve in any environment between z=0.8 and the present for galaxies with $M>4\times 10^{10}$ Msol

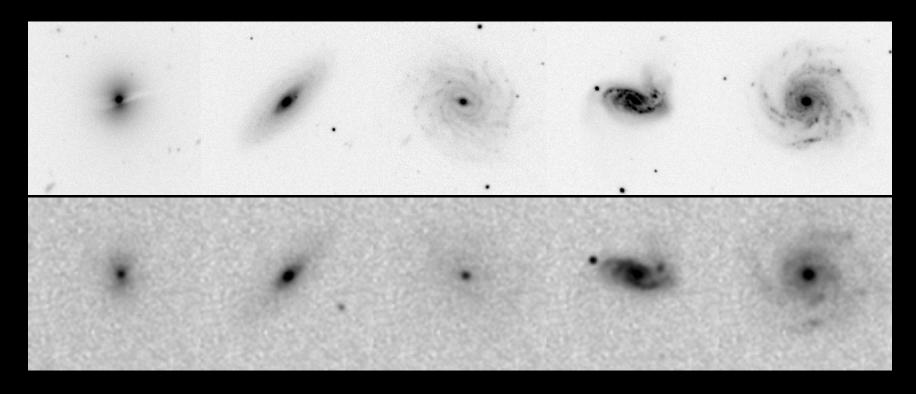
What WFC3 will bring...



L* - 2 L*;
$$H(AB) = 23 \dots 24$$
 at $z = 2.2$

4 orbits with F160W

What WFC3 will bring...



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