# Studying the Very High Redshift Universe with Gravitational Telescopes

### Johan Richard (Caltech)



Richard Ellis, Dan Stark (Caltech) Eiichi Egami (U. of Arizona) Jean-Paul Kneib (Marseille) Graham Smith (Birmingham)

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# <u>Outline</u>

- Motivation : towards the Dark Ages
  Use of Lensing clusters as Gravitational Telescopes
- Search for lensed dropouts with HST (ACS/NIMOS)
   Photometric selection and dropout candidates
   Reliability and redshift estimation
  - Spectroscopic follow-up
  - **Implications**



# Motivation : end of the Dark Ages

#### <u>WMAP</u> :

Reionization epoch :  $z \sim 10-12$ 

<u>QSOs</u>:
 I.G.M. fully reionized at z ~ 6

Get constraints on
 \* Nature (stars, AGNs)
 \* Physical properties
 \* Formation epoch
 of the sources responsible
 for reionizing the IGM.



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# Identification of high redshift galaxies

#### Selection with Broad Band Filters Lyman-Break Galaxies (LBG)

#### 2.5 < z < 4.5 Steidel et al (96 -> 05) <u>Ex: Bouwens et al. 2004, 2006, z~7-8</u> <u>Candidates at z~7-8 selected on HST/UDF field</u> zJH photometry

#### **Our current motivation is exploratory :**

- no clear spectroscopic confirmation of a galaxy at z > 7

- number counts of earliest galaxies
- contribution of stellar formation to reionization

#### **Techniques used:**

- Study of photometric dropouts
- Use of strong lensing by massive clusters: extend these techniques to faint luminosities



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#### Strongly-lensed Lyman- $\alpha$ emitters : Stark et al. 07



Systematic survey with NIRSPEC/Keck of the critical lines in 10 redshift clusters, produced 6 Lyman- $\alpha$  candidates at 8.5 < z < 10.4

 $\Rightarrow$  high density of low luminosity SF galaxies at z~10?

We would like to check this with another technique: lensed dropouts.

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### Beyond z~6 with Strong Gravitational Lensing

#### Kneib, Ellis, Santos, Richard 2004:

z ~ 6.8 I-dropout
 Confirmed with the well-constrained lensing model + photometric redshift

(No Ly- $\alpha$  emission detected)



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### Detection of the z~7 object with SPITZER





Detection of both images at λ> 4000 Å (Rest-frame) :
→ physical properties indicate earlier star-formation
(Egami, Kneib, Rieke, Ellis, Richard et al. 2005, ApJL)

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### **Searching for Lensed Dropouts with HST/Spitzer**



Richard et al (in prep.)

Critical line: infinite mag

Mag.  $> \times 6$ 

- 6 well-constrained clusters with deep IRAC imaging (Egami & Rieke)
- 11 NICMOS pointings in 6 lensing clusters (4 orbits J/F110W, 5 orbits H/F160W)
- ACS/F850LP imaging of all clusters
- K-band ground based imaging with Keck/NIRC + Subaru/MOIRCS

### **Combining ACS, NICMOS & Spitzer**

#### MS1358: 5 $\sigma$ limit: J<sub>AB</sub>=26.7, H<sub>AB</sub>=26.7



#### Importance of foreground removal

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### Lensed z-band dropouts (z~7-8)

• NIR colors evolution with redshift for E, Sb, Sc and Im templates + reddening

• Including two variations of color-color criteria, with z-J > 0.8 or z-J > 1.2



10 candidate z-drops in the 6 clusters surveyed with H ~ 26 - 26.8
z ~ 1-2 red galaxies expected to be main contaminants

# **Reliability and redshift estimation (1)**

• False positive detections : tests on "noise image"

Noise image created by subtracting in pairs NICMOS science frames.



• Estimation: in the magnitude range of the dropouts, we expect ~ 10 % spurious detections (i.e. 1 out of the 10 dropouts)

# Reliability and redshift estimation (2)

• Effects of post-overexposures on the detector : measurements from our own data and the archive: no effect

• Low-mass stars : L and T dwarfs are expected to contaminate the survey. Predictions: 1 star in entire survey.

 Photometric redshifts
 Contamination by lower z galaxies: estimation of 25 % from P(z)



 $\Rightarrow$  statistically expect 6 out of 10 dropouts likely to be at high z

# Search for multiple images

 Counter-images predictions from lensing model

2 candidates with possibly "merging" images

Proof of Method: we do see z~2 multiple sources...

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### **Properties of stacked SED for the z-dropouts**





Best fit  $z_{ph}=7.35 + -0.12$ (J - H)<sub>AB</sub> ~ 0.05 UV spectral slope best fit:  $f_{\lambda}=\lambda^{-\beta}$  $\beta \sim 2.6-2.8$ 

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# Keck/NIRSPEC spectroscopic follow-up

Optimization to follow-up both a candidate and its predicted counter-image



- NIRSPEC slit : 0.76" x 42"
- Follow-up in the Z band (6.8<z<8.3 for Lyman-alpha)
- 3 to 4 hours on 7 candidates

Sensitivity to lyman α flux: 3 e-18 erg/s/cm2 in 50% of the spectrum, or W>25 Angstr.

No detection of Ly- $\alpha$  : possibly different distribution of equivalent widths between  $z \sim 6.0$  (Stanway 07) and  $z\sim7.5$ 

### Implied Luminosity Function z~7.5

No significant overlap between UDF and lensed survey



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#### **Implications for Reionization from Lensed Dropouts**



• Even if a few are real, suggests significant contribution to reionization from low luminosity galaxies

• Consistent with picture revealed by lensed Lyα emitters (Stark et al 2007)

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### Galaxies at z ~ 10



- Selection criterion : J-H > 1.8
- 2 good candidate J-drops all 6 clusters, each with  $H_{AB}$ ~25.5 25.7
- SFR ~ 0.1 1  $M_{\odot}$  yr<sup>-1</sup> (unlensed)
- 1 of them not detected in K-band with MOIRCS/Subaru

# Summary

• Evidence for early star formation beyond  $z \sim 7$  is seen in current surveys: this occurred either in extincted objects or, more likely, in low luminosity systems

• Strong lensing surveys are finding an abundant population of faint dropouts at z  $\sim$  7-10, with SFR <  $\sim$  1  $M_{\odot}$  yr^-1

• Spectroscopic follow-up shows no detection of Ly-a : possibly different distribution of equivalent widths between  $z \sim 6.0$  and  $z \sim 7.5$ 

• These programs, and upcoming dedicated instruments such as WFC3 will give a first glimpse of the Universe at z > 7, and more effectively plan ambitious programs with EMIR/GTC, JWST and ELTs